

Photometer and lidar measurements in Spitsbergen

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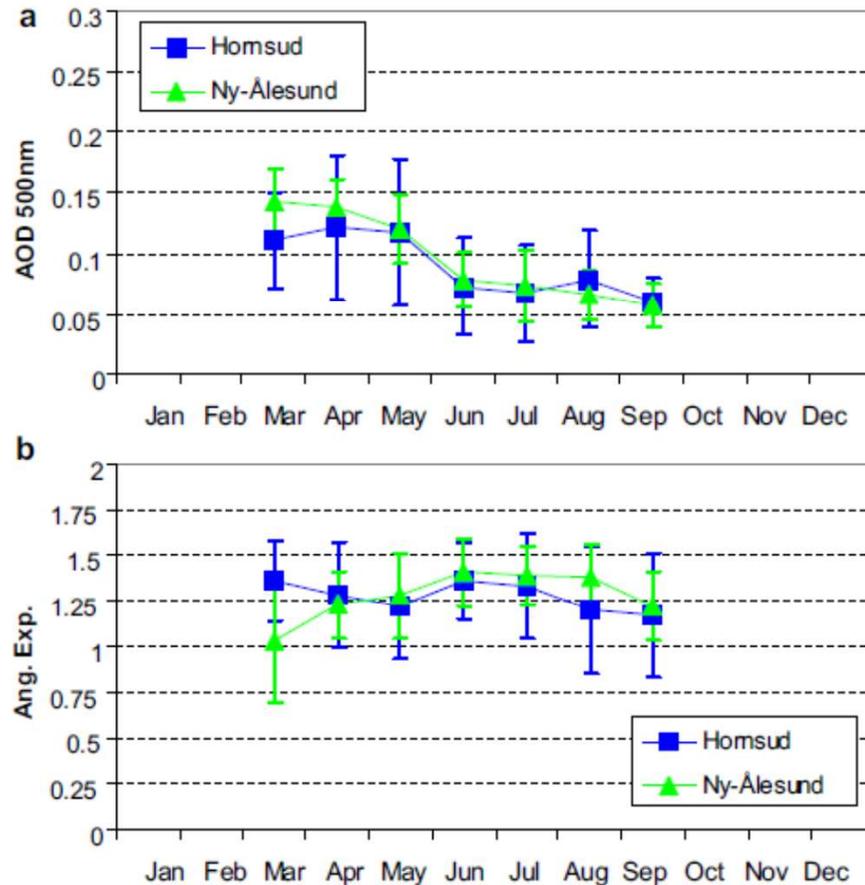
And:

Sandra Graßl, sandra.grassl@awi.de

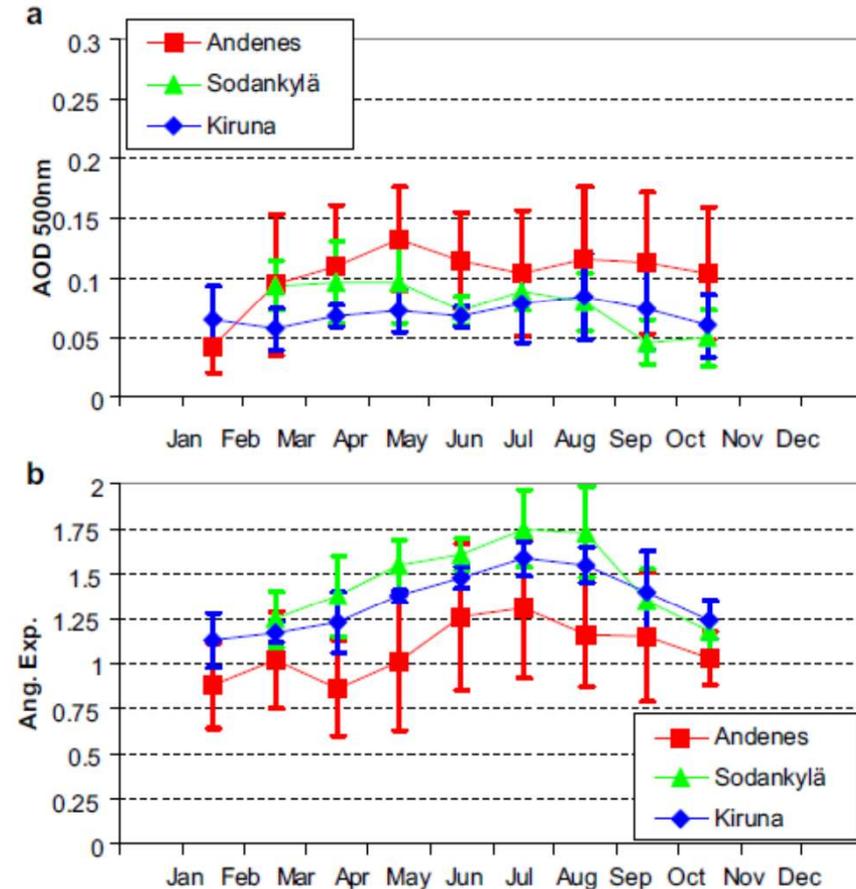


Typical AOD values from Toledano 2012 Atmos. Environm.

Spitsbergen



Scandinavia

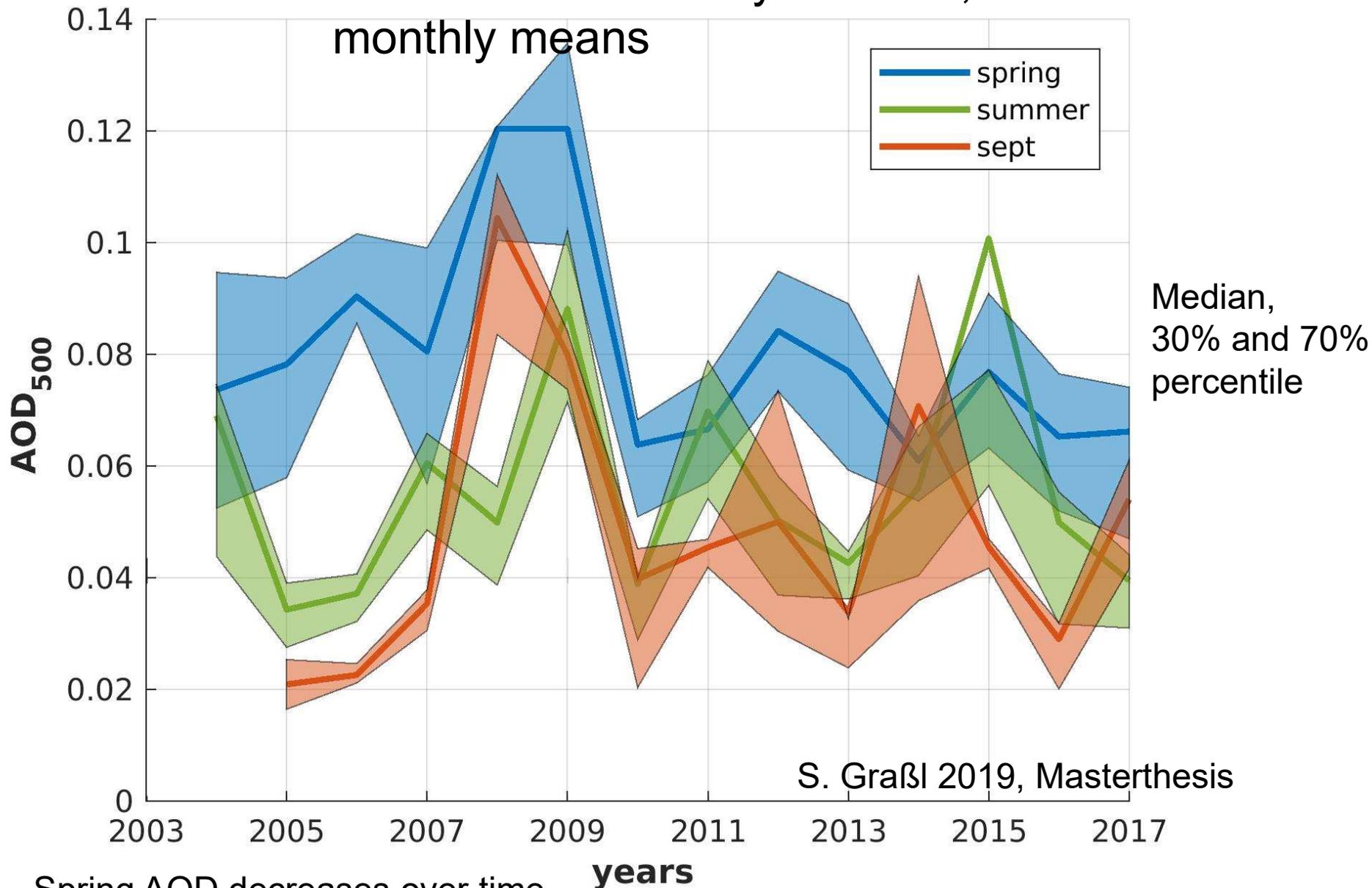


Spring: Arctic-AOD > N-European-AOD
 No Haze in Scandinavia
 No „easy“ direct pollution transport from Europe

Contrary: Eckhardt 2003 (Flextra, CO Tracer) „NAO + facilitates transport into Arctic“

Aerosol may have different pollution pathways than trace gases!

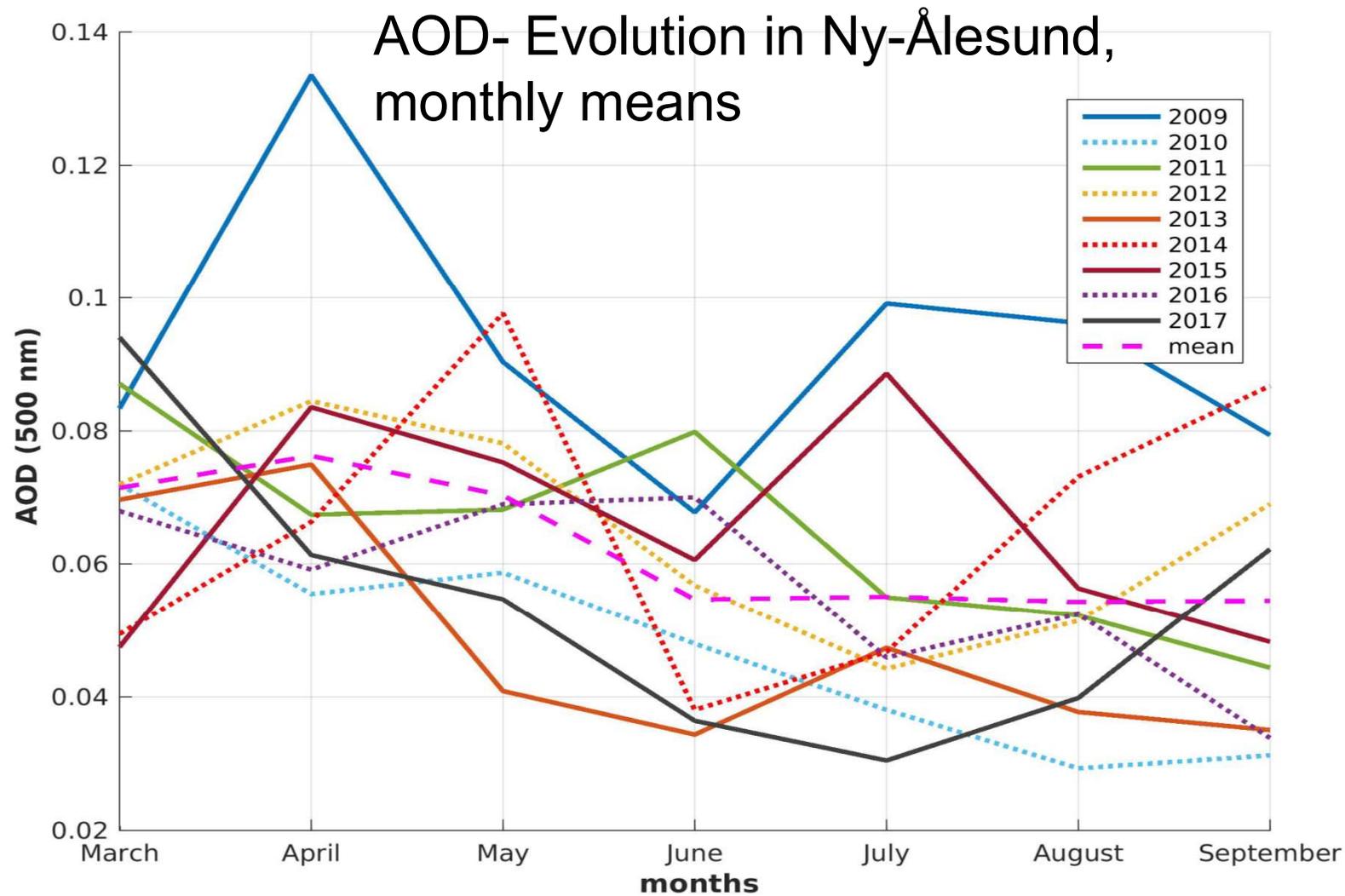
AOD- Evolution in Ny-Ålesund, monthly means



Spring AOD decreases over time

→ annual run of AOD becomes flatter

2009 was last polluted year Generally high variability

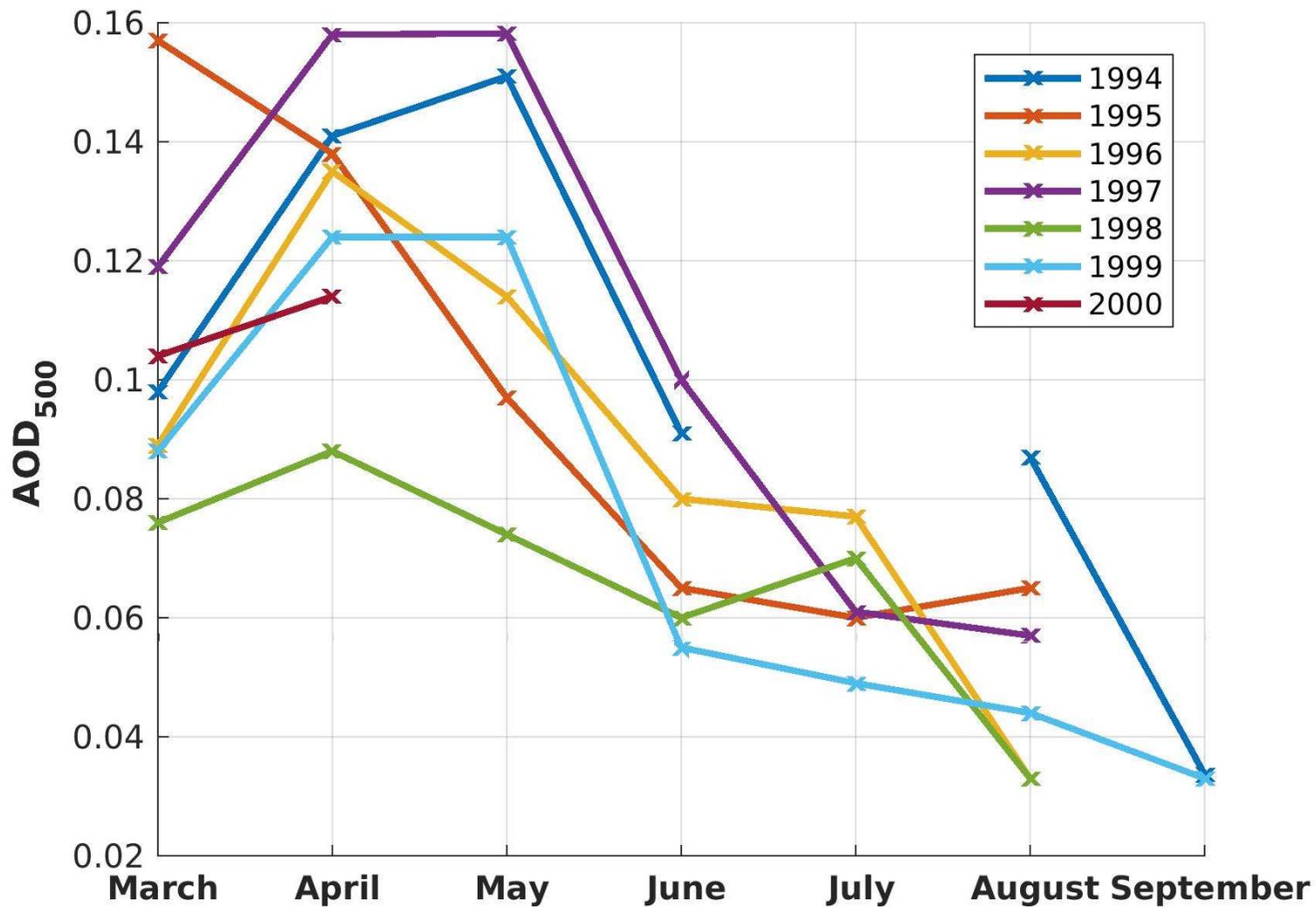


S. Graßl 2019, Masterthesis

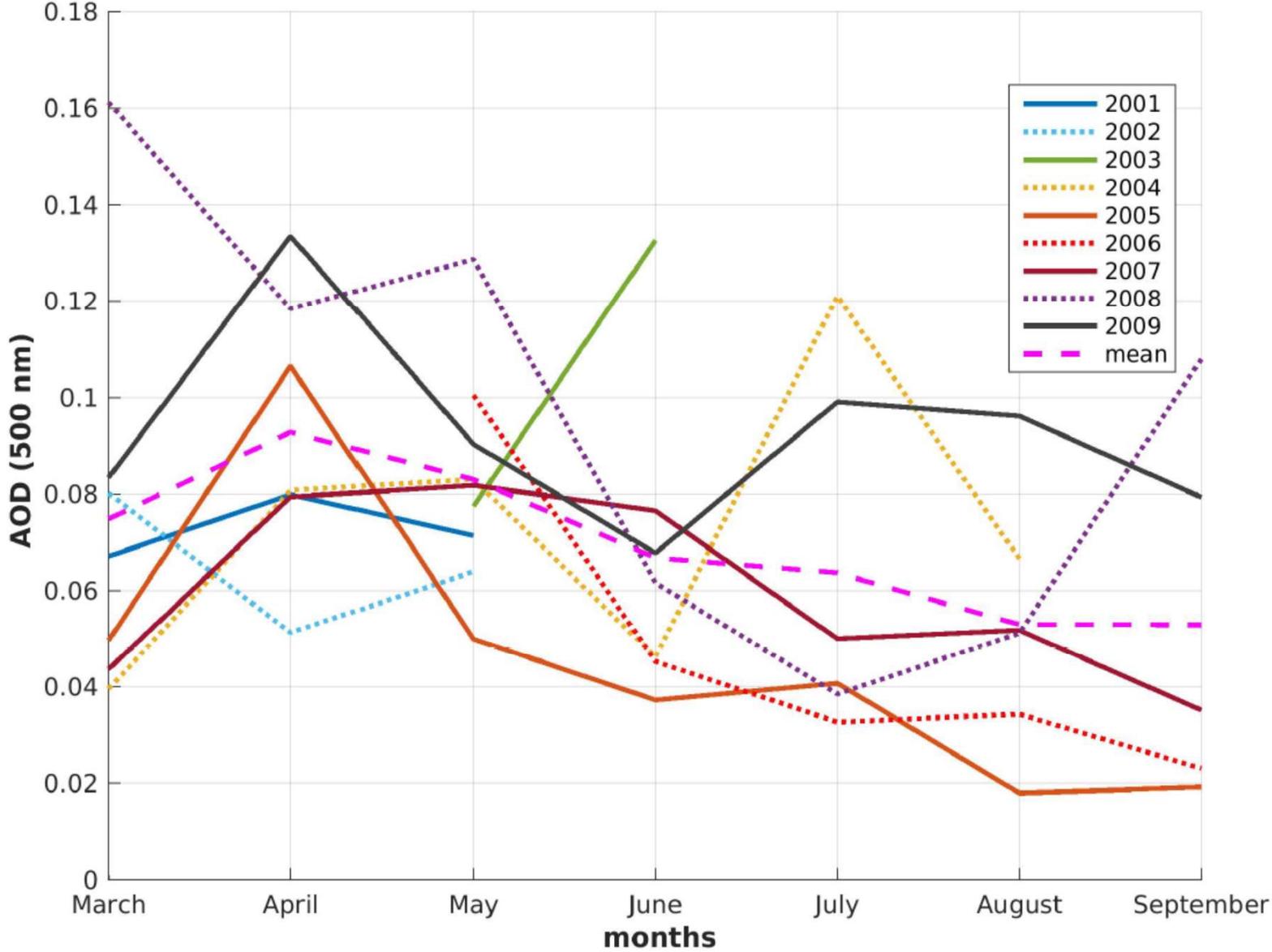
Spring AOD decreases over time
 → annual run of AOD becomes flatter
 2009 was last polluted year
 Generally high variability

Jul- Sep 2009: Mt Sarychev

Old date from Herber 2002:
More Haze and longer Haze periods, (still in May!)

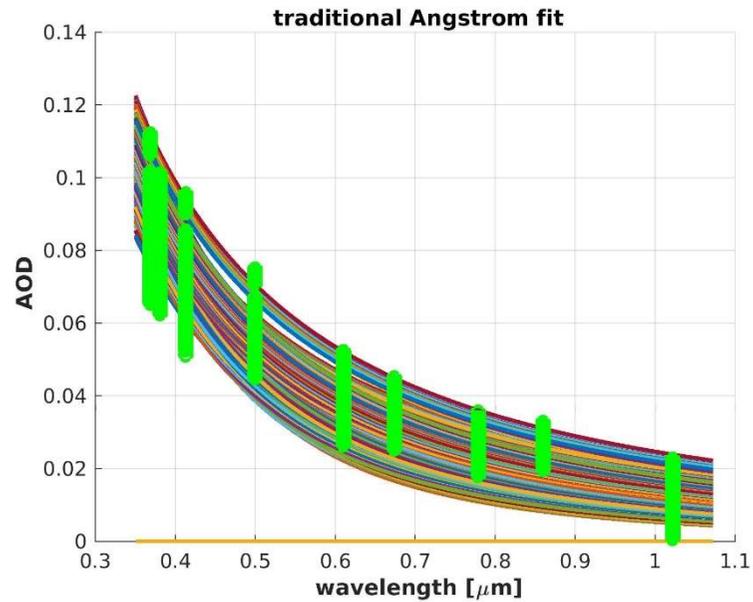


And the years in between, AOD is shrinking but with high variability



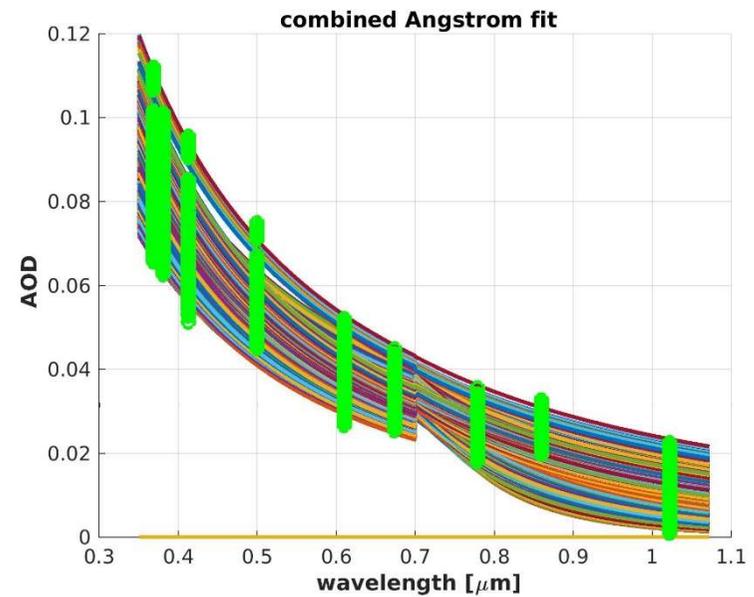
New approach of photometer evaluation:

1) Information content: (many data from 2013)



Traditional Angstrom exponent:

$$\tau = C \lambda^a$$



Same data set but separate fits for $\lambda >$ or $<$ 700nm

Angstrom exponent for data Ny-Alesund wavelength dependent

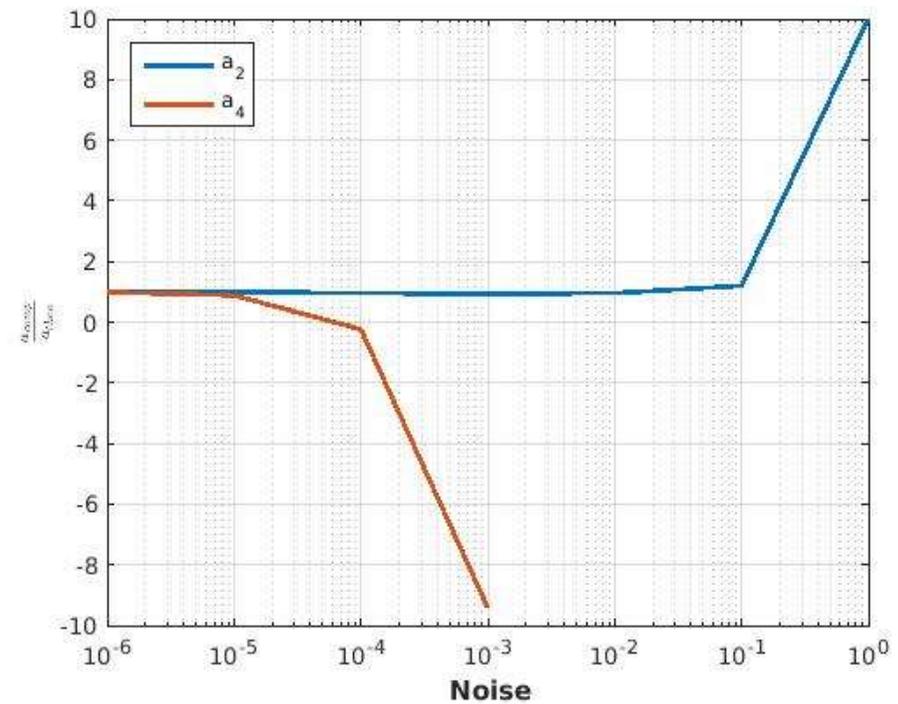
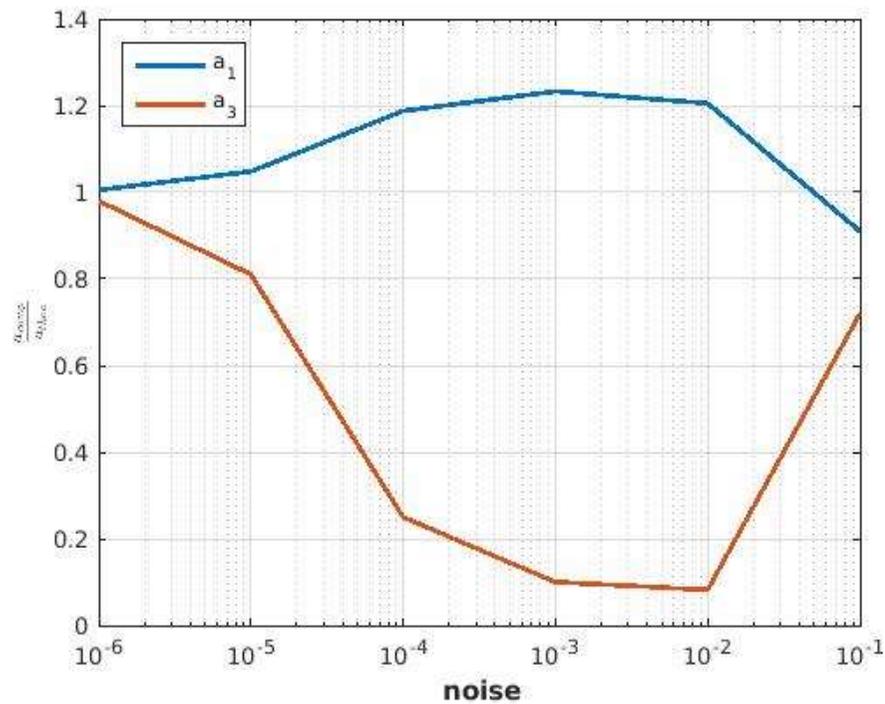
In an ideal world one could imagine:

$$\tau = a_1 \lambda^{a_2} + a_3 \lambda^{a_4}$$

coarse - fine- mode

Method: invent $a_1, a_2, a_3, a_4 \Rightarrow \tau_{\text{theo}}$ add noise

Use Levenberg-Marquardt to retrieve a_1, a_2, a_3, a_4 back



Hence: this approach will only work if $\frac{\Delta\tau}{\tau} = O(10^{-5})$

Hence we are looking for an easy approach which contains more information than the traditional Angstroem and chose:

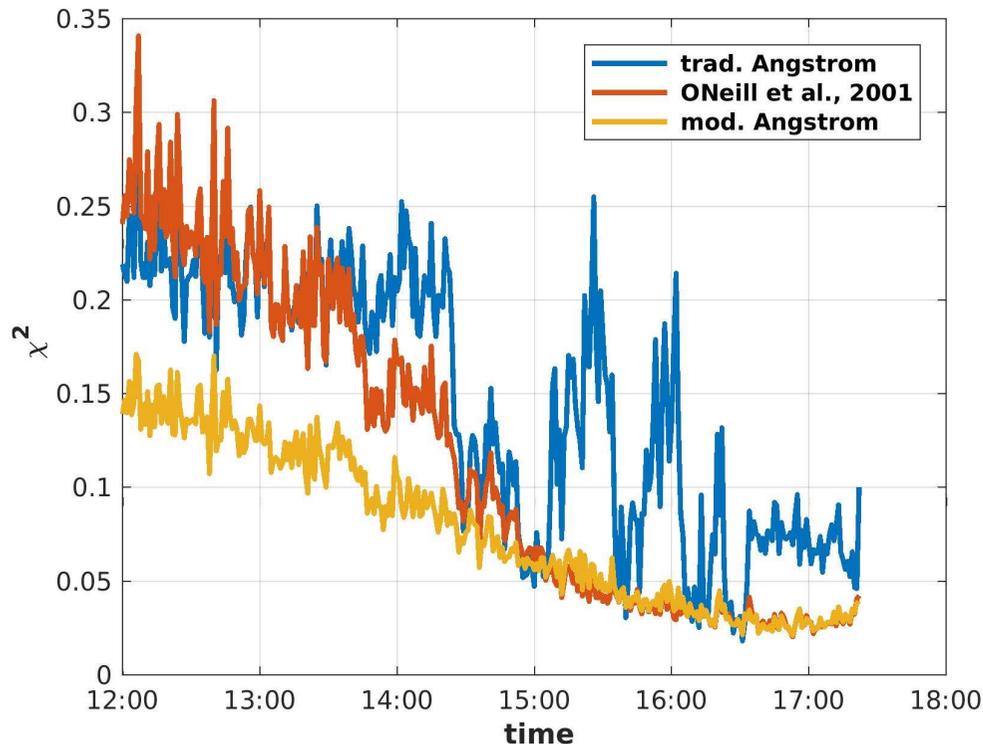
$$\tau = C \lambda^{-\alpha + \beta \cdot \lambda}$$

α : „modified Angstroem“ β : „spectral slope“

α : AE (λ) = 0

β : d/d λ (AE)

Again we use the LM to retrieve C, α , β from the measured $\tau(\lambda)$
 Fitting possible up to $O(10^{-1})$



$$\tau = C \lambda^{-AE}$$

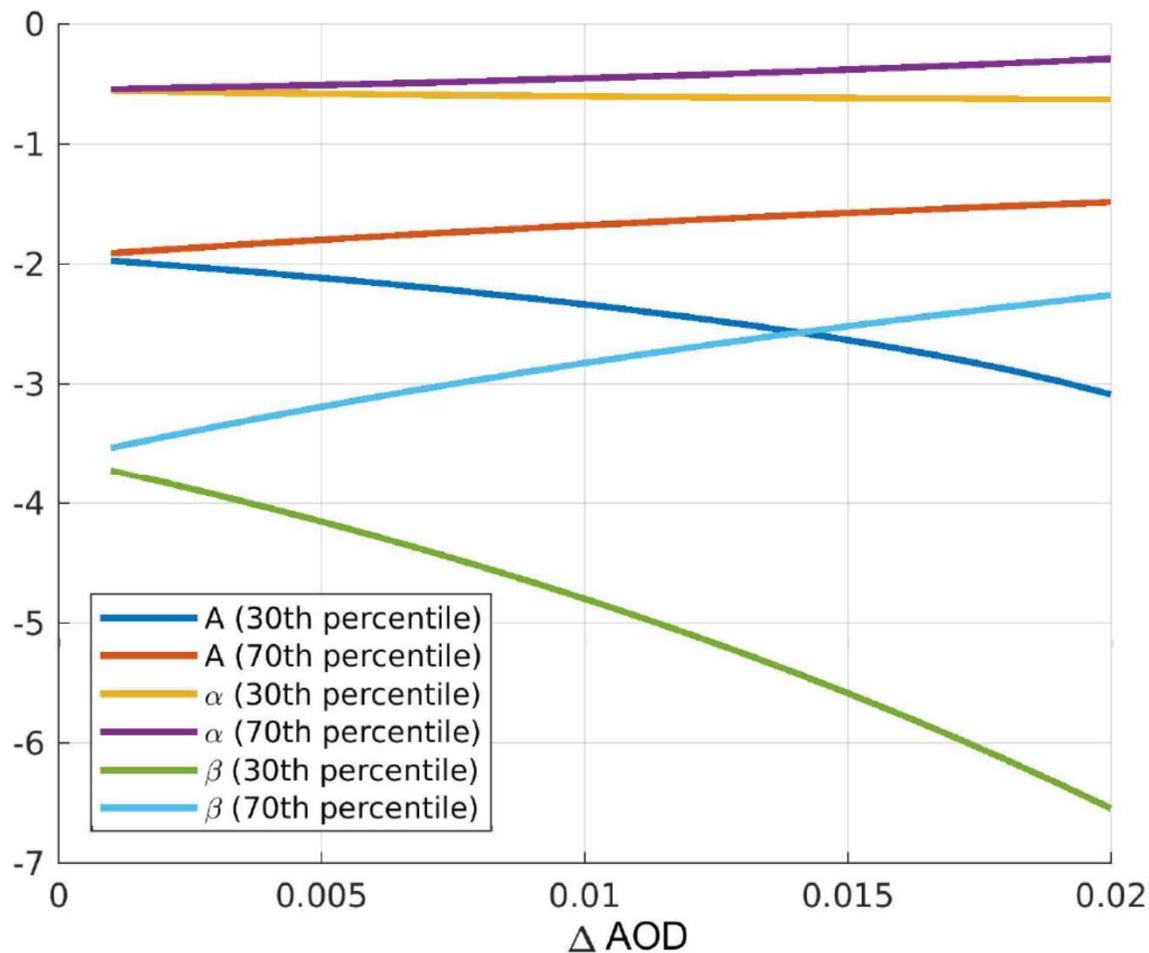
$$\log \tau = C + \log \lambda + (\log \lambda)^2$$

$$\tau = C \lambda^{-\alpha + \beta \cdot \lambda}$$

Information content of photometer data:

Approach:

- 1) Chose arbitrary $\tau(\lambda)$
- 2) Add 100 noise realisations with given ΔAOD
- 3) Retrieve for each noise realisation A , α , β

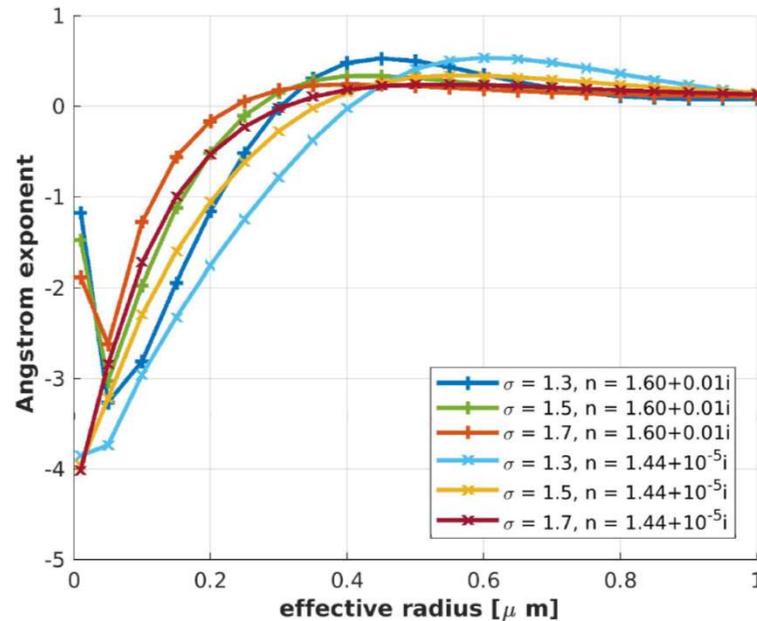


FWHM spread:
(for $\Delta AOD = 0.01$)

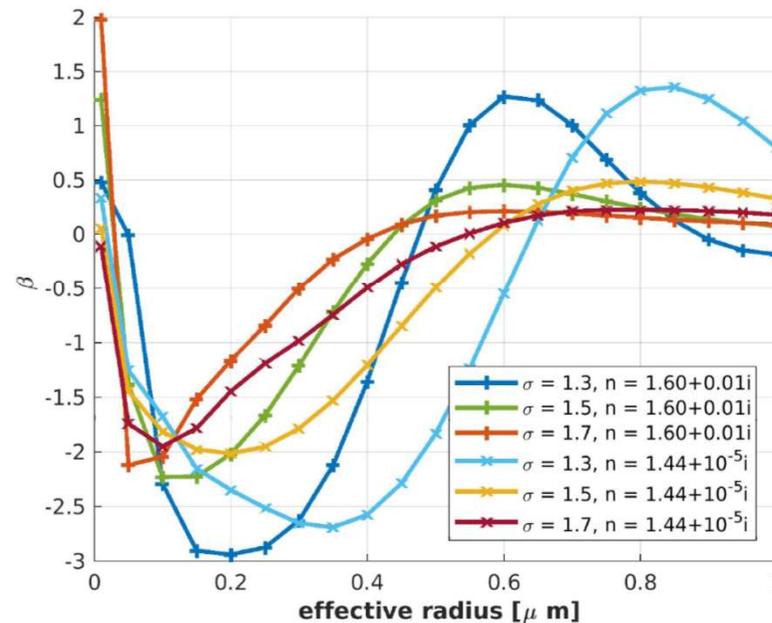
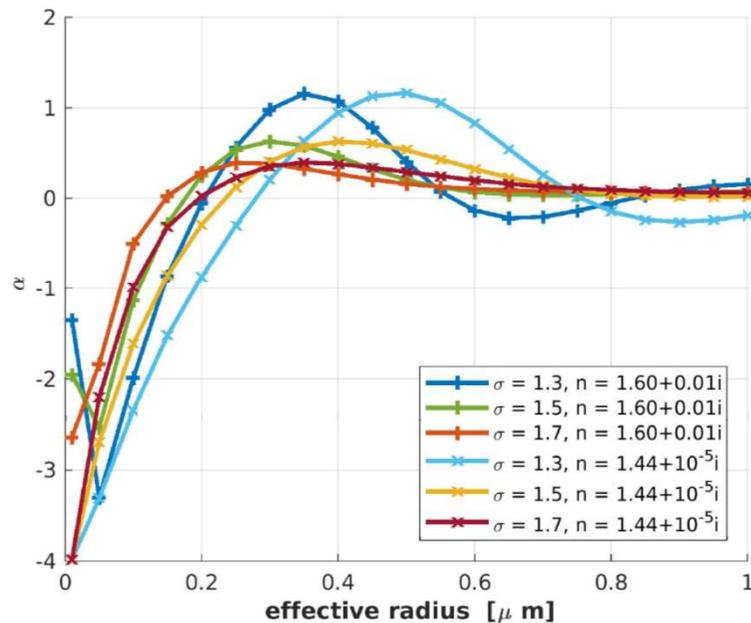
$A = 0.14$
 $\alpha = 0.66$
 $\beta = 1.96$

Information content behind photometer data:

Invert log-normal distributions (σ , n , r_{eff})
 Mie: \rightarrow retrieve
 A , α , β



Above 0.4 μm grey approximation approaches



What do α and β tell us?

Both are intensive quantities, so they can be plotted against each other independent of aerosol concentration

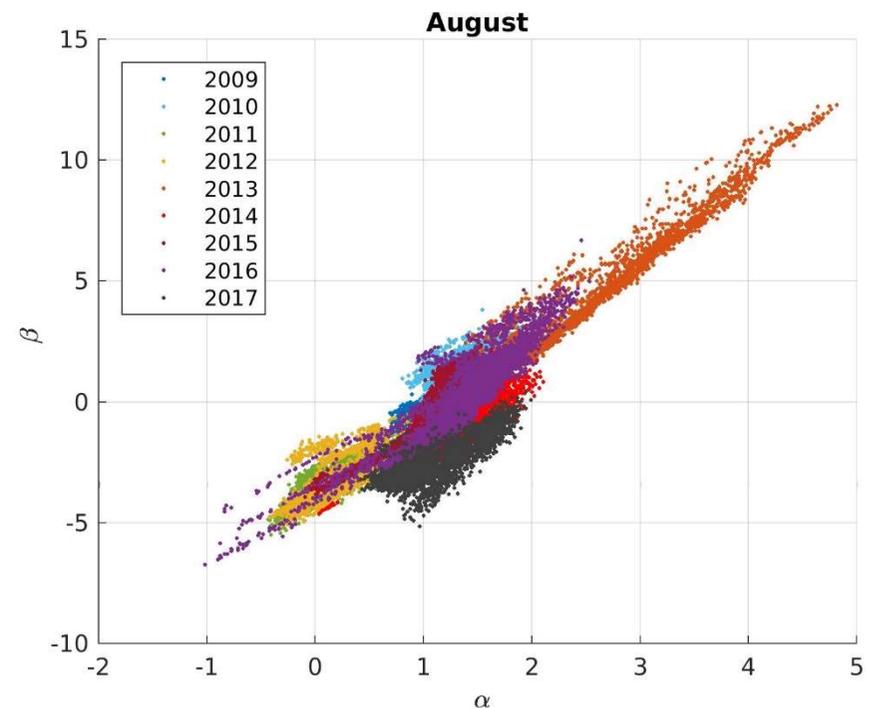
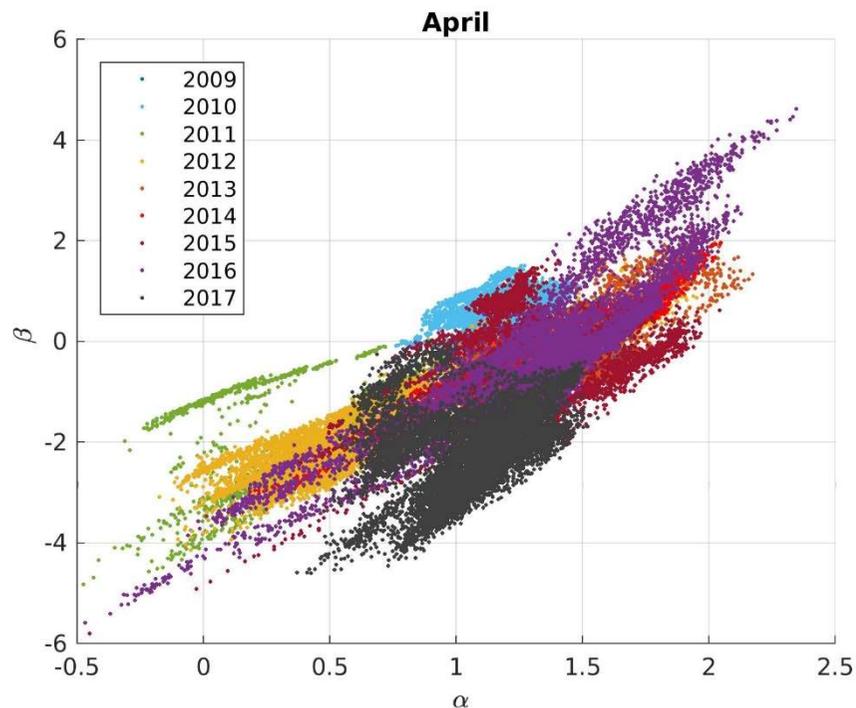
Recall: $\tau = C \lambda^{-a+b \cdot \lambda} = C \lambda^{-Exp}$

Hence: $\beta < 0$: AE stronger negative for IR (Expon. for IR closer to -4)

$\beta = 0$: traditional Angstrom law good

$\beta > 0$, requires small α

Often Exponent smaller for IR than for UV



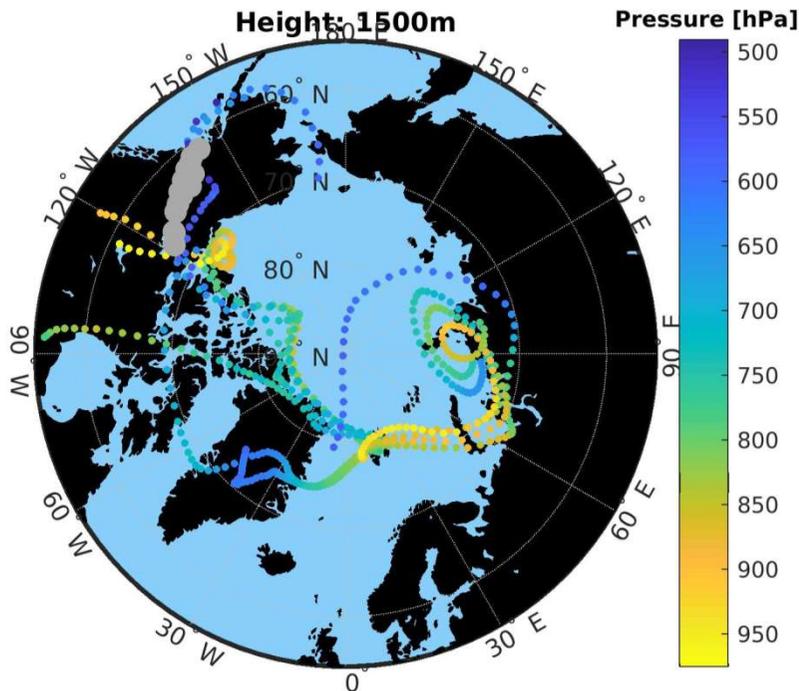
Non-uniqueness of size estimation from A, α , β

refractive index $n_1 = 1.60 + 0.01i$			
σ	A	α	β
1.3	$r_{eff,1} = 0.01\mu m$ $r_{eff,2} = 0.18\mu m$	$r_{eff} = 0.13\mu m$	$r_{eff,1} = 0.11\mu m$ $r_{eff,2} = 0.34\mu m$
1.5	$r_{eff} = 0.14\mu m$	$r_{eff} = 0.09\mu m$	$r_{eff,1} = 0.09\mu m$ $r_{eff,2} = 0.20\mu m$
1.7	$r_{eff} = 0.10\mu m$	$r_{eff} = 0.07\mu m$	$r_{eff,1} = 0.11\mu m$ $r_{eff,2} = 0.34\mu m$
refractive index $n_2 = 1.44 + 10^{-5}i$			
σ	A	α	β
1.3	$r_{eff} = 0.16\mu m$	$r_{eff} = 0.10\mu m$	$r_{eff,1} = 0.09\mu m$ $r_{eff,2} = 0.16\mu m$
1.5	$r_{eff} = 0.21\mu m$	$r_{eff} = 0.14\mu m$	$r_{eff,1} = 0.06\mu m$ $r_{eff,2} = 0.35\mu m$
1.7	$r_{eff} = 0.28\mu m$	$r_{eff} = 0.19\mu m$	$r_{eff,1} = 0.05\mu m$ $r_{eff,2} = 0.16\mu m$

Assuming A= -1, α =-1, β =-1.5 (typical for Ny-Ålesund) different solutions are possible

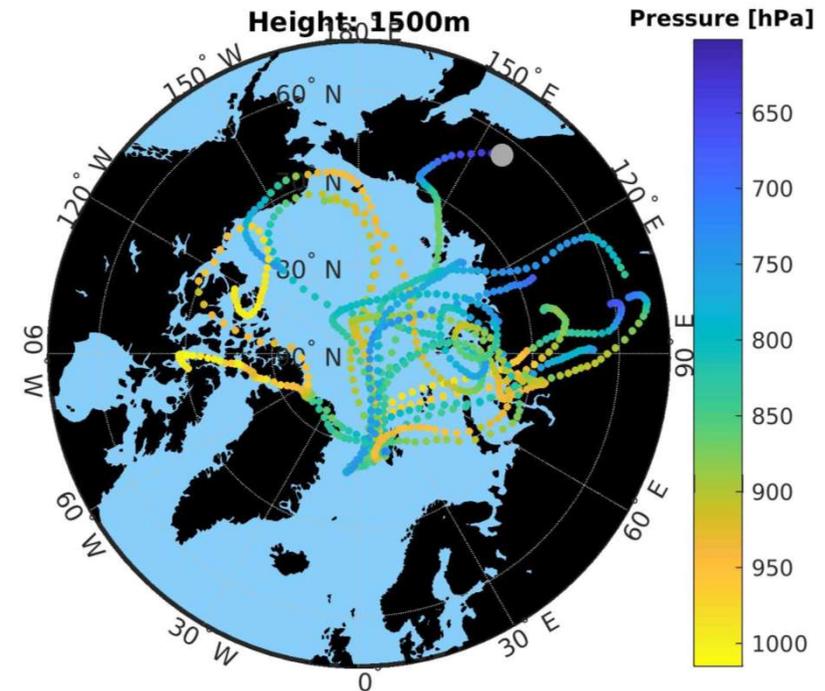
Open questions: Pollution pathways

Graßl, 2019: Flextra with ERA-interim



Low AOD

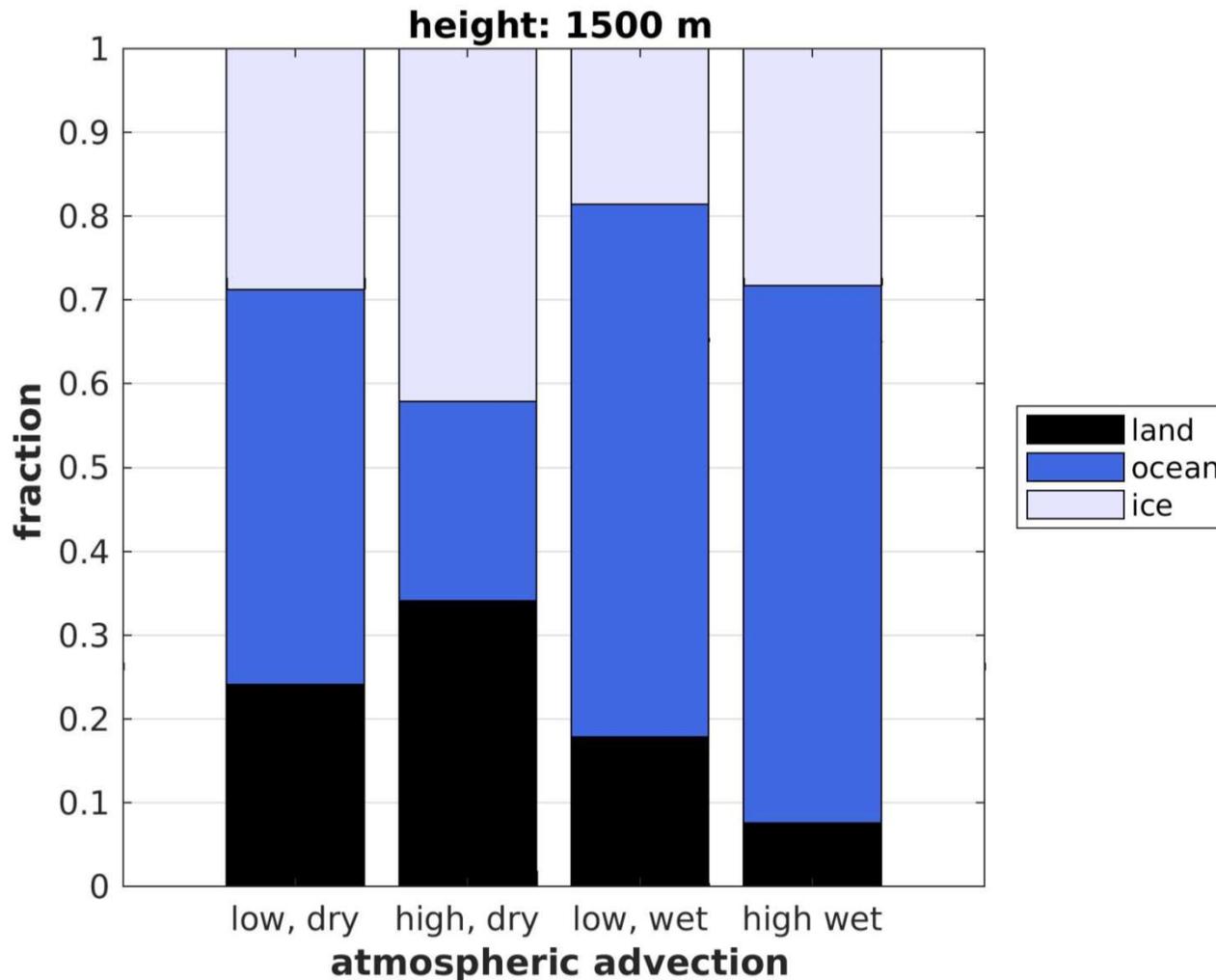
(April 2013)



high AOD

5 days trajectories too short
Reanalysis products show large differences
Slightly higher AOD from Siberia

Sea ice as reduced sinks?



High aerosol load due to sources and sinks

Sea ice: dry, stable BL
less vertical mixing,
longer aer. life-time

Best conditions for aerosol transport:

Air over source regions in BL with enough wind speed

Ascend of the air (higher wind speed, 5 days, less precipitation)

Advection over sea ice

FLEXTRA 5 days (with photometer) Aprils 2013-2016

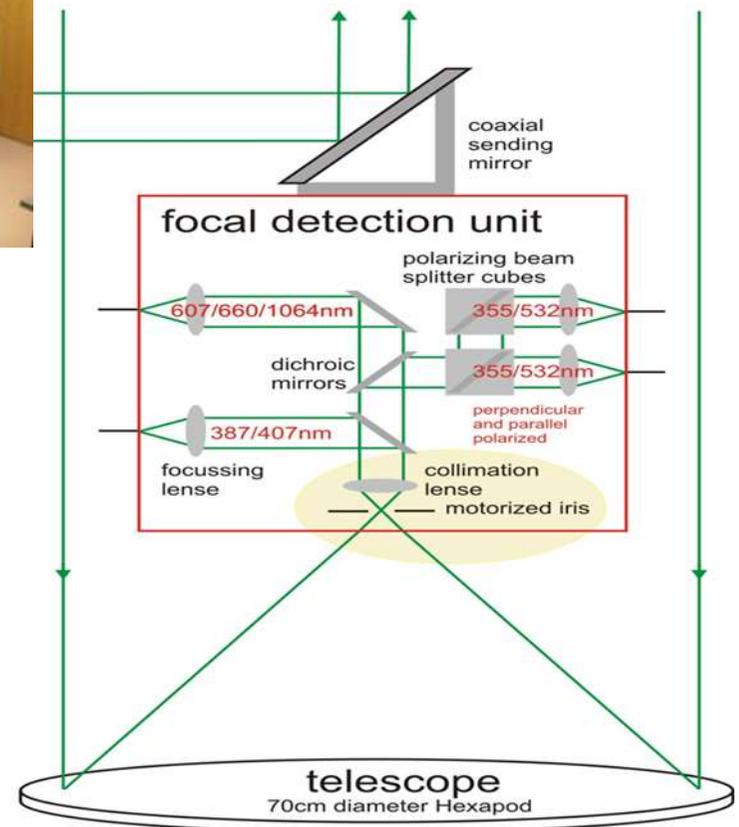
MOSAIC: coordinated observations with surrounding stations needed

KARL: Koldewey Aerosol Raman Lidar

Backscatter (β) @ 355nm, 532nm, 1064nm
 Extinktion (α) @ 355nm, 532nm
 Depolarisation (δ) @ 355nm, 532nm
 Water vapor (m_r) @ 407nm, 660nm



Spectra 290 /50 Laser (10W / colour)
 70cm mirror
 Fov: 1 4 mrad
 Licel transients, Hamamatsu PMTs
 Overlapp > 700m
 Tropo- & stratosphere



What does an aerosol lidar deliver:

extensive quantities (dependent on aerosol number concentration):

backscatter (concentration, size, shape, refractive index)

extinction (concentration, size, shape, refractive index) !

Intensive quantities (not dependent on aerosol number concentration)

depolarisation $\delta = \frac{\beta_{\perp}}{\beta_{=}}$ (shape) [dipole moment]

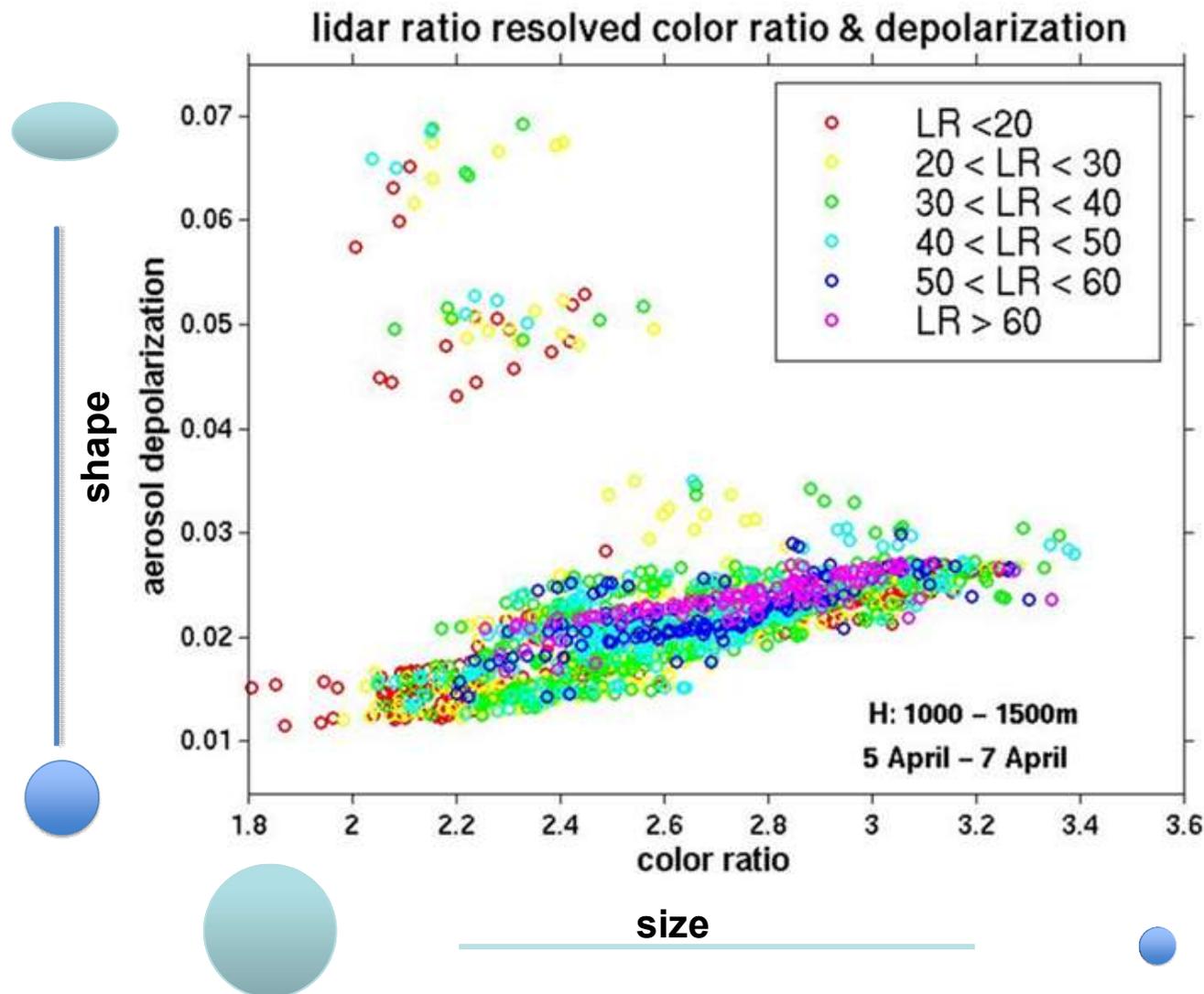
colour ratio $CR = \frac{\beta_{\lambda_1}}{\beta_{\lambda_2}}$ (size) [$\beta \sim \lambda^{\text{\AA}}$ $-4 < \text{\AA} < 0$]

lidar ratio $LR(\lambda) = \frac{\alpha^{aer}}{\beta^{aer}}$ (index of refraction, size, shape)

Knowledge of δ , CR, LR allows a robust classification of aerosol type (dust, smoke, sea salt, cirrus...)

→ it's about getting the intensive quantities!

Mixing state of aerosol:



Sort aerosol for size and shape: still very inhomogeneous LR:

Chemistry unrelated to size and shape

On scale 30m/ 10min no individual soot, sulphate, crust ... particles

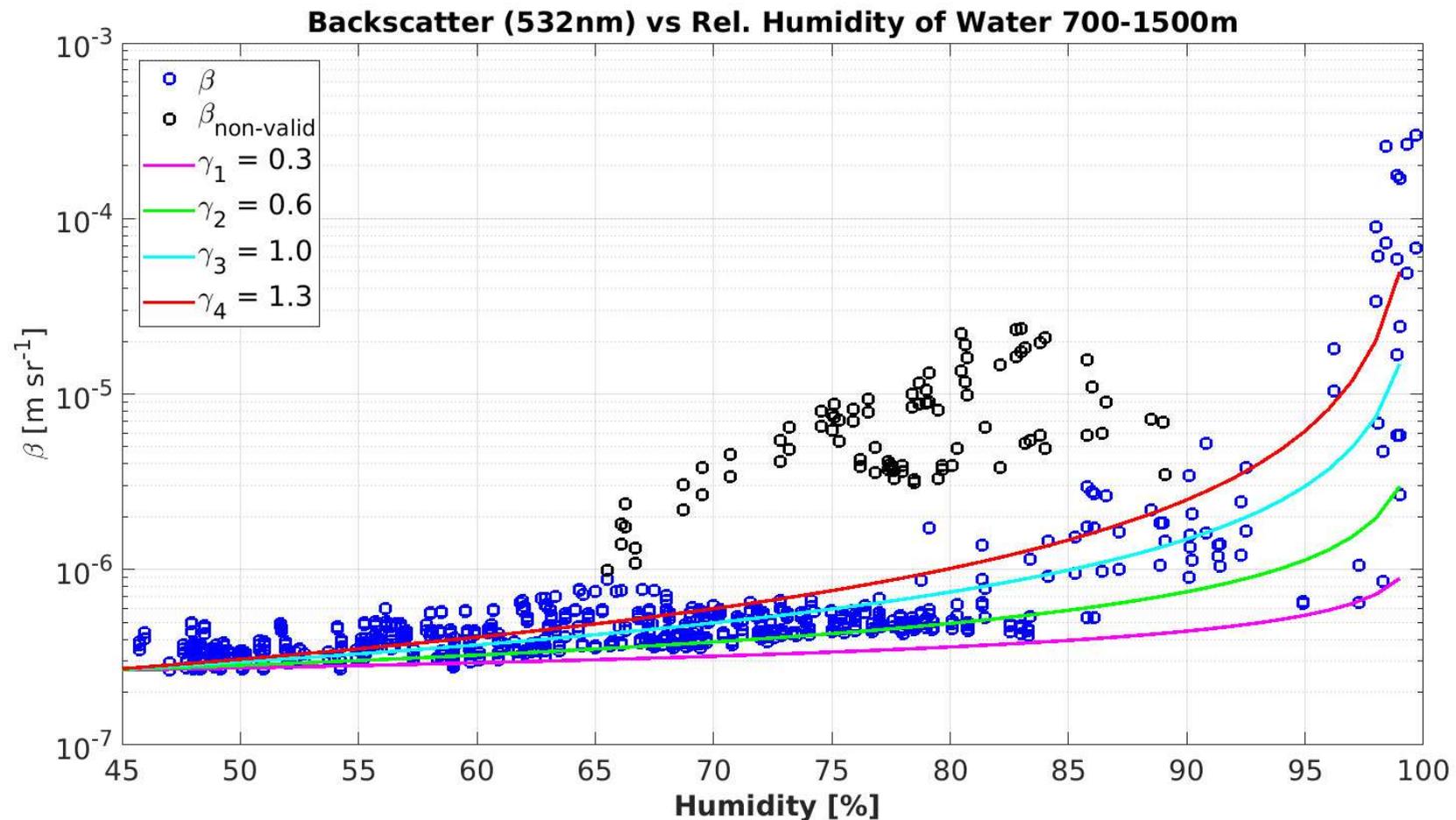
Color ratio, depol. ratio both intensive quantities

Lidar and contemporary radiosonde: hygroscopic growth?

In-situ define scattering enhancement factor $f(\text{rh}) = (1-\text{rh})^{-\gamma}$

Question: apply this to β (instead of σ)?

Assumption: all lidar data in a given time / height should belong to „same event“



A theory is short,
concise and
complete and is
believed by nobody
except of its
inventor.



Observational data
are noisy, strange
and incomplete and
are believed by
everybody except
of the one who
measured them.

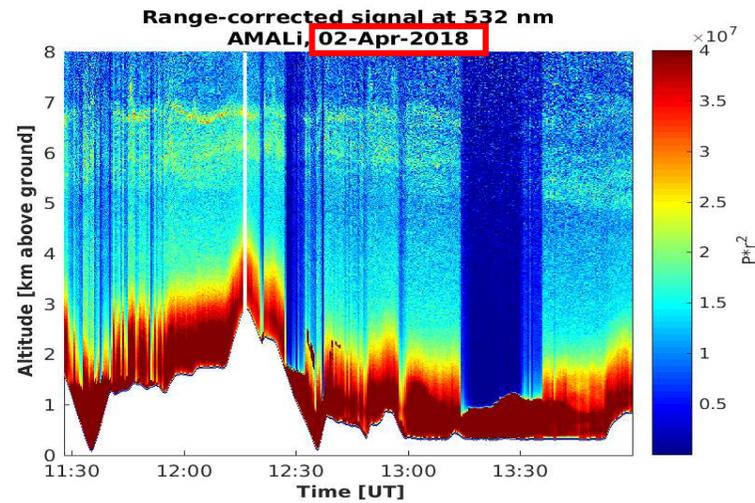
Picture:
Loriot 1923 - 2011

Thank you for your attention!

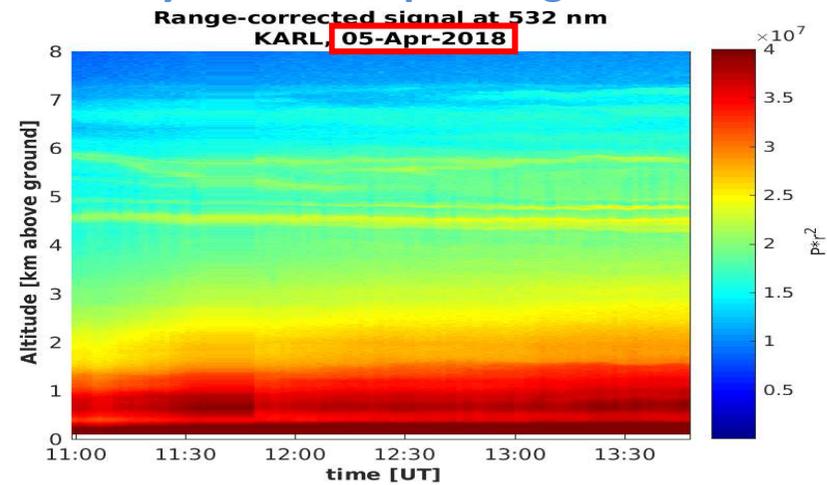
AC3 and PAMARCMiP 2018:



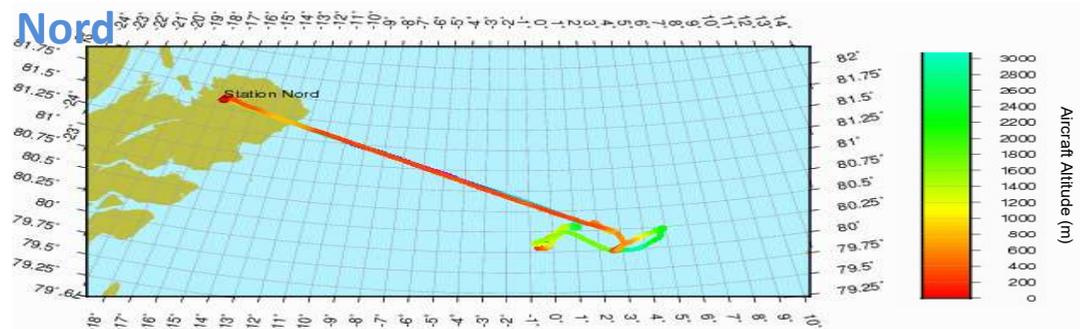
Persistent layer of aerosol in 5-7km



Ny-Ålesund, Spitsbergen



Polar5 flight-track towards Station Nord



Compare remote sensing to in-situ

Calculate radiative forcing

Open questions:

2. Does remote sensing overestimates extinction?

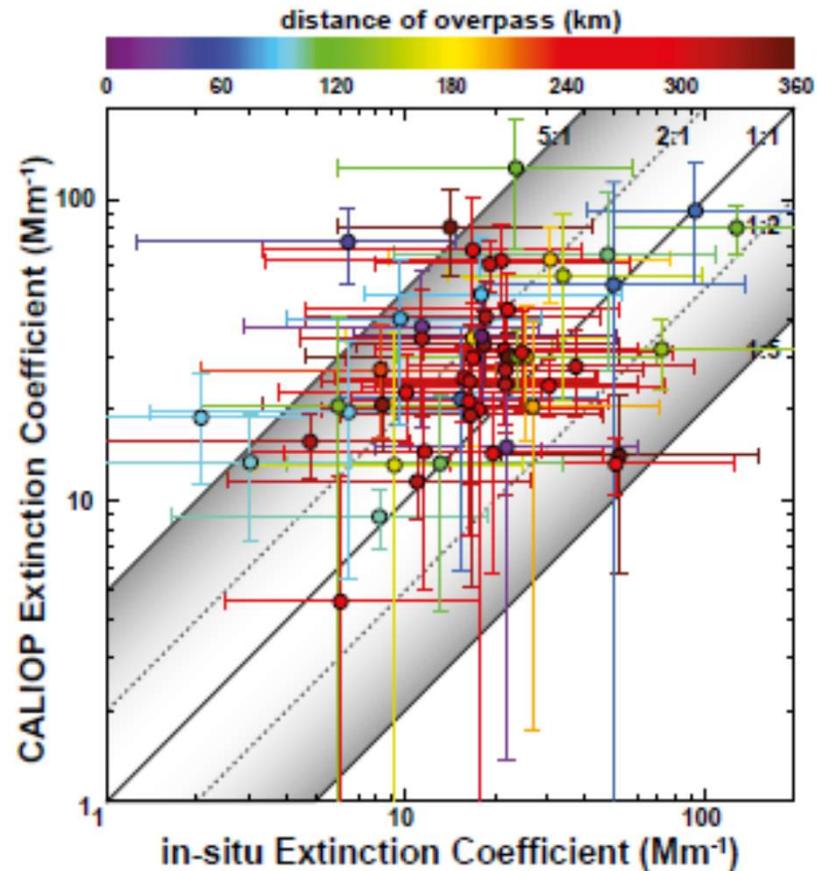
Tesche et al. 2014 ACP:

Calipso_extinction > in-situ
(Zeppelin station)

(what was NOT published in)
Lisok, 2016 Atm. Environm:

KARL_extinction > in-situ
(Gruvebadet station)

And extinction at ground, 1km,
2km altitude not correlated
Deviations also at $rh = 50\%$



Needs to be clarified during MOSAiC:
Less orography!