

CO₂ reconstructions and carbon cycle in the past (20 Myr)

ESSReS-L9

Earth System Science: a combined data-modelling paleoperspective

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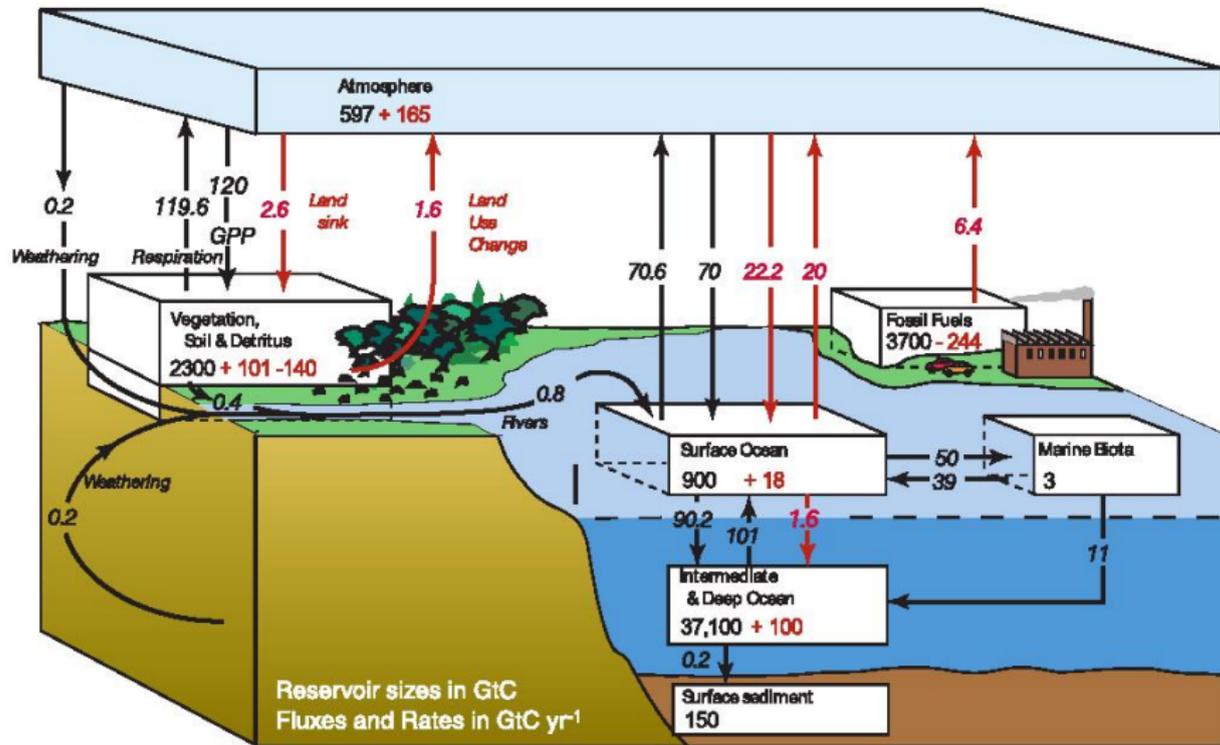
04 May 2011

- 1 Basics on the Carbon Cycle
- 2 CO₂ reconstructions
 - $\delta^{11}\text{B}$
 - B/Ca
 - Alkenones, $\delta^{13}\text{C}_{\text{org}}$
 - Stomata
 - Validation of different approaches
 - Greenhouse Effect
- 3 Processes
 - The Faint young sun Paradox
 - CO₂ outgassing
 - Weathering
- 4 Summary

Outline

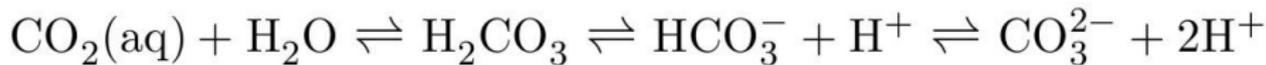
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C Pools and C fluxes



CO₂ in Seawater

CO₂ in seawater reacts with water and dissociates immediately after:



Only the part of CO₂, which get dissolved after Henry's Law can exchange with the atmosphere.

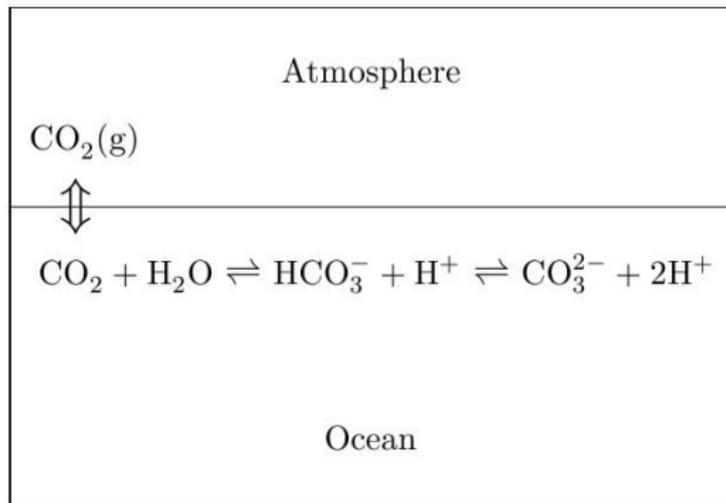
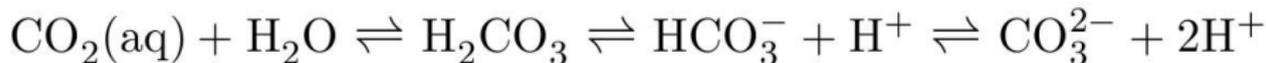
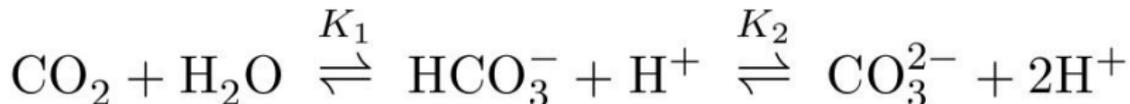


Figure 1.1.1: Schematic illustration of the carbonate system in the ocean. CO₂ is exchanged between atmosphere and ocean via equilibration of CO₂(g) and dissolved CO₂. Dissolved CO₂ is part of the carbonate system in seawater that includes bicarbonate, HCO₃⁻, and carbonate ion, CO₃²⁻.

Chemical System in Equilibrium



[H₂CO₃] is negligible and the equation reduced to



Dissolved Inorganic Carbon — DIC

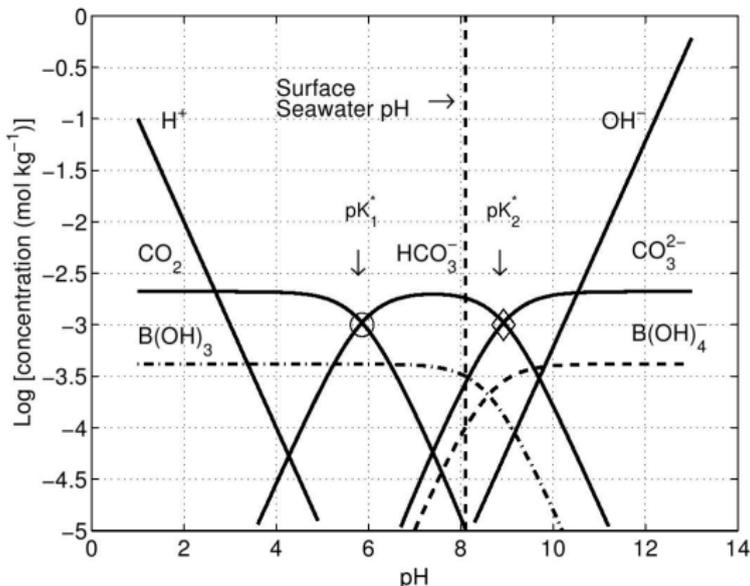
$$\text{DIC} \equiv \Sigma\text{CO}_2 = [\text{CO}_2] + [\text{HCO}_3^-] + [\text{CO}_3^{2-}]$$

DIC, ΣCO_2 also sometimes called PCO_2

Equilibrium constants:

$$K_1^*, K_2^* = f(\text{temperature } T, \text{ salinity } S, \text{ pressure } P).$$

Bjerrum Plot



Present day conditions and $S = 35$, $T = 25^\circ \text{C}$:

$[\text{CO}_2] = 10 \mu\text{mol kg}^{-1}$; $[\text{HCO}_3^-] = 1818 \mu\text{mol kg}^{-1}$; $[\text{CO}_3^{2-}] = 272 \mu\text{mol kg}^{-1}$

$[\text{CO}_2] : [\text{HCO}_3^-] : [\text{CO}_3^{2-}] \sim 1\% : 90\% : 10\%$

Zeebe & Wolf-Gladrow 2001

Total Alkalinity

Total Alkalinity (TA or ALK) is the excess of proton (H^+ ion) acceptors over proton donators (with respect to a zero level of protons).

Or even simpler:

Proton acceptor: negative charged ion

Proton donator: H^+ or ion/molecule that can spend one H^+ ion

Roughly:

$$TA \sim 1 \times [HCO_3^-] + 2 \times [CO_3^{2-}]$$

also called **carbonate alkalinity**

Or in detail:

$$TA = 1 \times [HCO_3^-] + 2 \times [CO_3^{2-}] + [B(OH)_4^-] + [OH^-] - [H^+] + \text{minors}$$

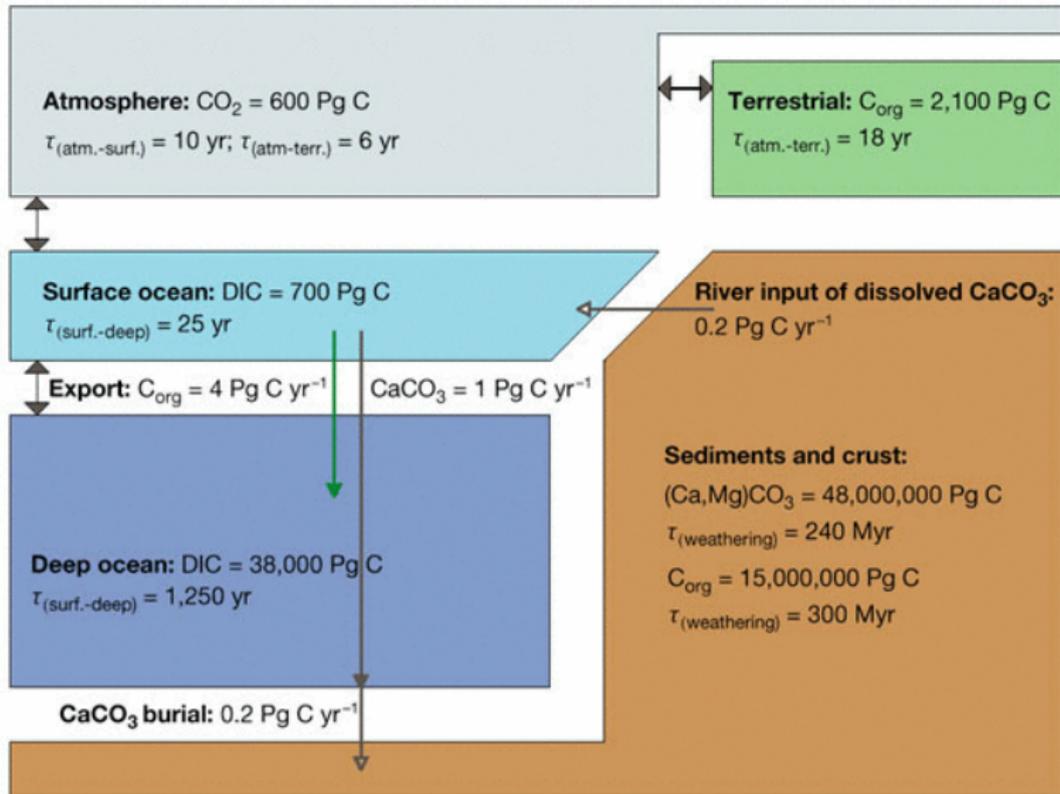
Carbonate System

Total Alkalinity and DIC are conservative quantities, meaning, their concentrations are unaffected by changes in pH , pressure, temperature, or salinity

CO_2 , HCO_3^- , or CO_3^{2-} are not conservative!

With two variables (out of DIC, TA, CO_2 , HCO_3^- , CO_3^{2-} , pH) together with T, S, P the carbonate system is fully described, the other four quantities can be calculated out of them.

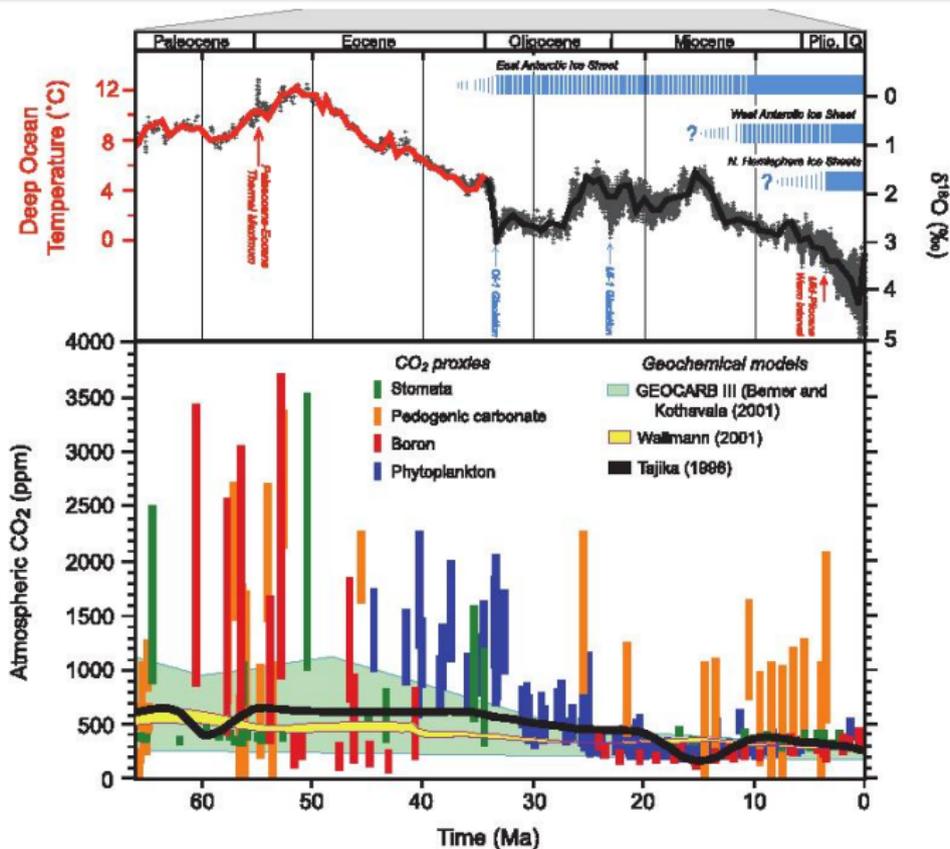
C Pools and C fluxes



Sigman and Boyle 2000 N

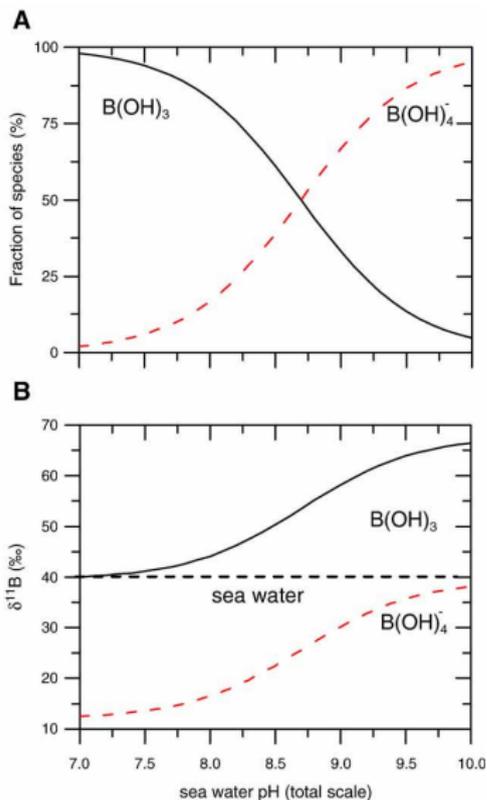
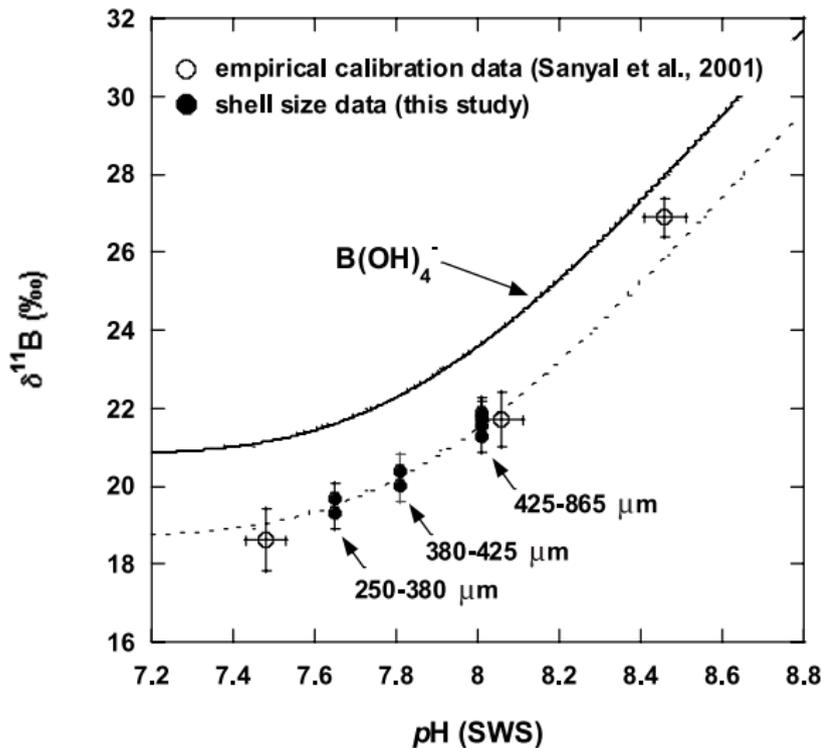
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CO₂ Reconstructions, 65,000,000 yr (IPCC 2007)

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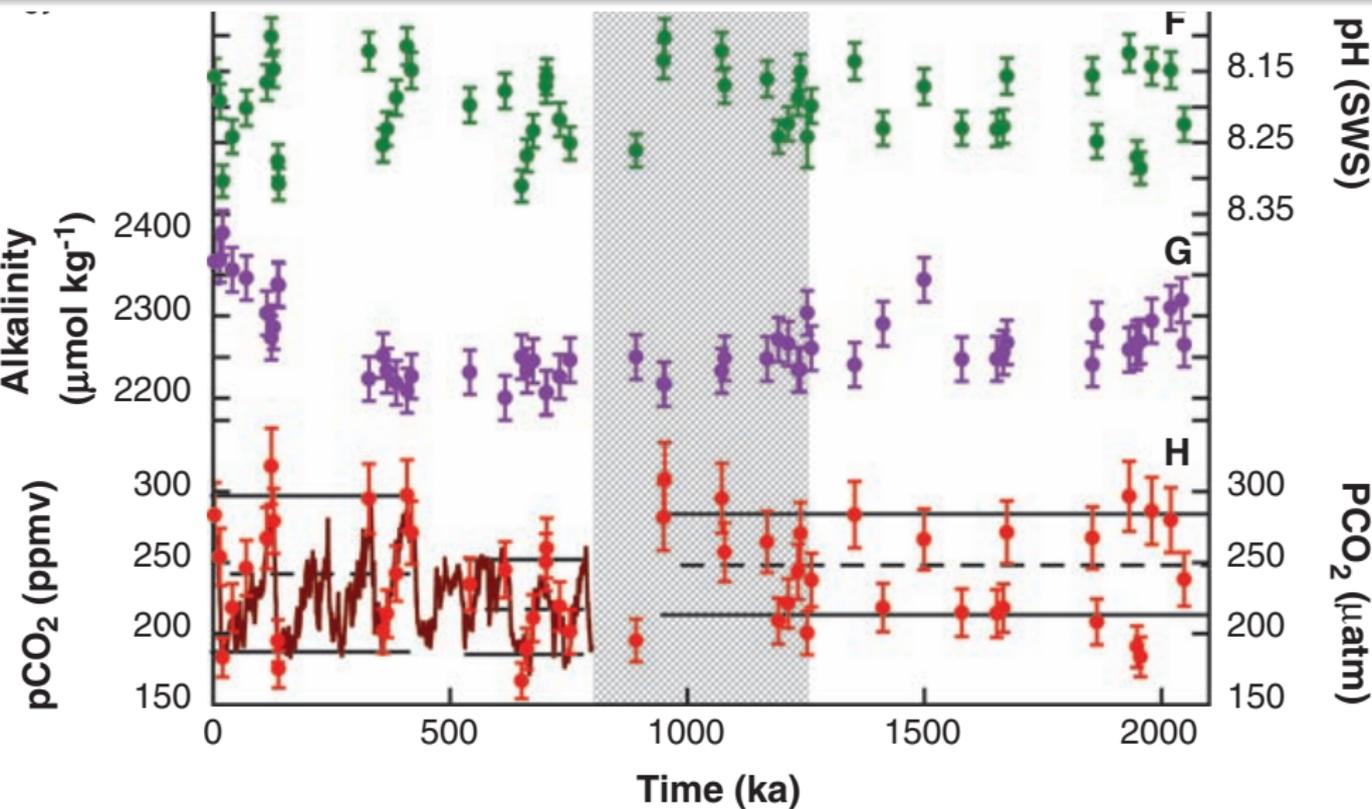
$\delta^{11}\text{B}$, pH— $\delta^{11}\text{B}$, pH—B*G. sacculifer*: shell size effect

Yu et al., 2010 EPSL; Hönisch 2004, P

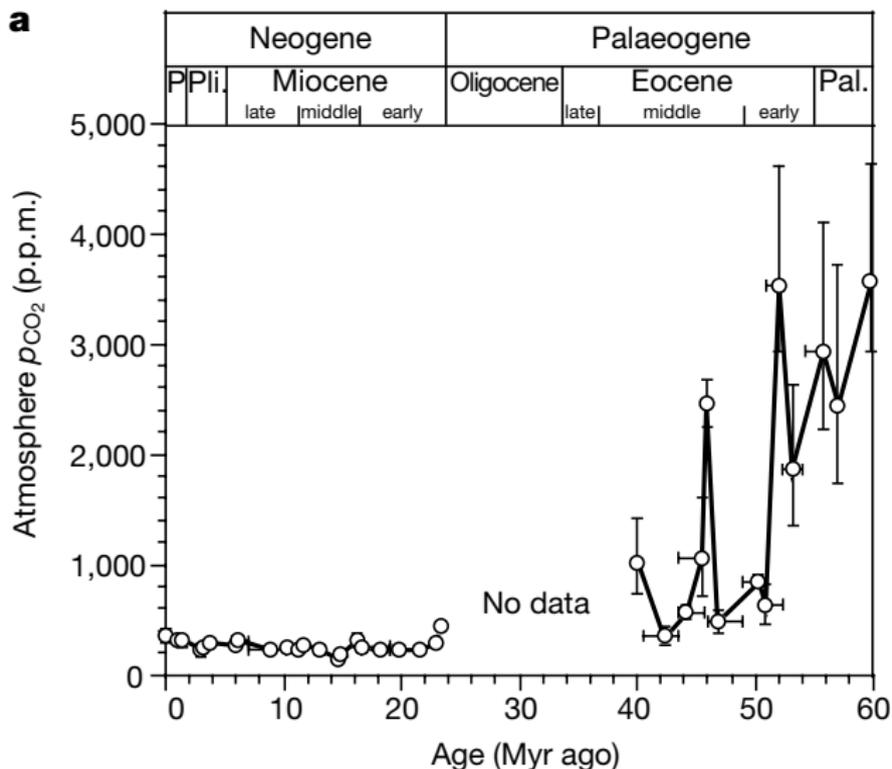
$\delta^{11}\text{B}$, boron isotopes

General approach:

- Calculate surface water pH out of $\delta^{11}\text{B}$.
- Determine independently another parameter of the carbonate system (CO_2 , HCO_3^- , CO_3^{2-} , pH, DIC, alkalinity), mostly alkalinity is estimated.
- Surface water $p\text{CO}_2$ can be calculated out of pH and 2nd parameter.
- Under the assumption that surface water $p\text{CO}_2$ and atmospheric $p\text{CO}_2$ stays (and stayed so in the past) in equilibrium this surface water $p\text{CO}_2$ is a proxy for atmospheric $p\text{CO}_2$.
- **Advantage:** Based on well understood marine chemistry
- **Disadvantage:** 2nd parameter needed, atm-surf-equilibrium might have changed over time, seems to work only for mono-specific selections

$\delta^{11}\text{B}$ example I, single species, last 2 Myr

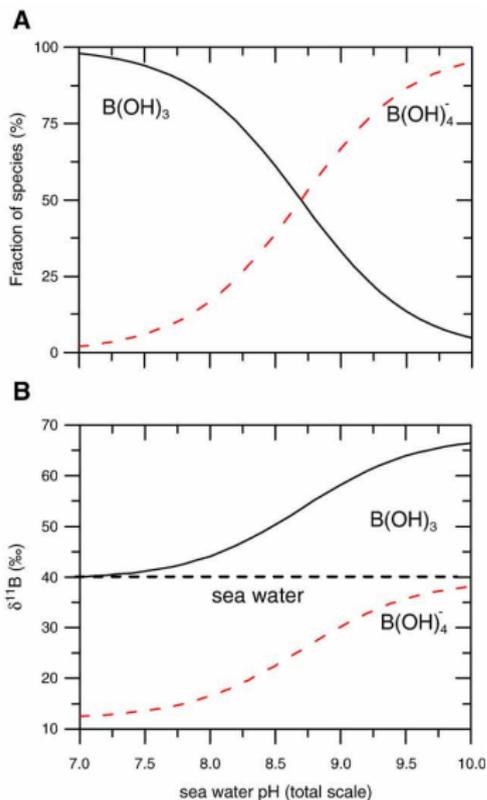
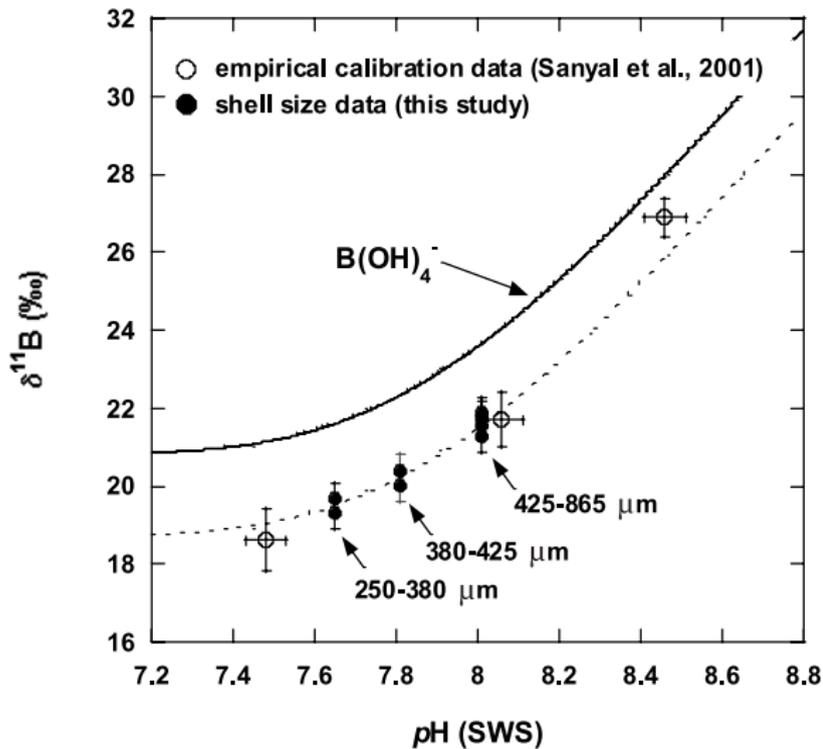
Hönisch et al 2009, S

$\delta^{11}\text{B}$ example II, multi-species, last 60 Myr

Pearson and Palmer 2000 N

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$\delta^{11}\text{B}$, pH— $\delta^{11}\text{B}$, pH—B*G. sacculifer*: shell size effect

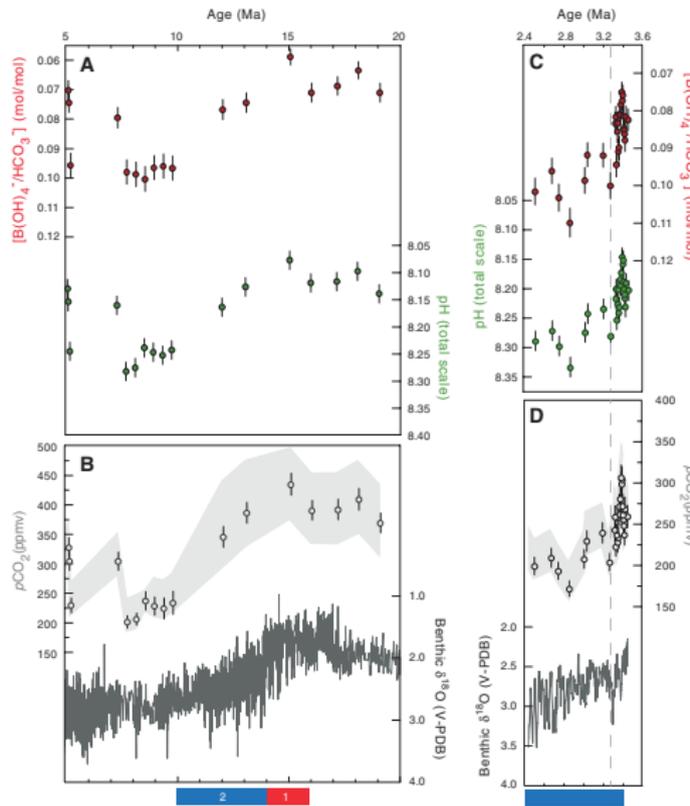
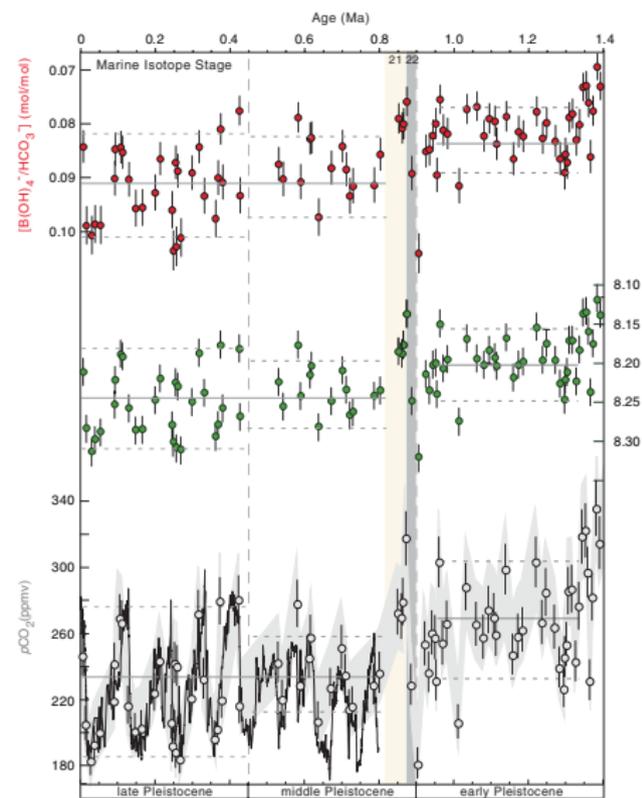
Yu et al., 2010 EPSL; Hönisch 2004, P

B/Ca

General approach:

- Planktic foraminiferal B/Ca ratios = f (seawater borate/bicarbonate ratios [B(OH)₄⁻/HCO₃⁻] = f(pH).
- similar to the $\delta^{11}\text{B}$ approach.
- **Advantage:** Based on well understood marine chemistry
- **Disadvantage:** 2nd parameter needed, atm-surf-equilibrium might have changed over time.

B/Ca example I, last 20 Myr



Tripathi et al 2009, S

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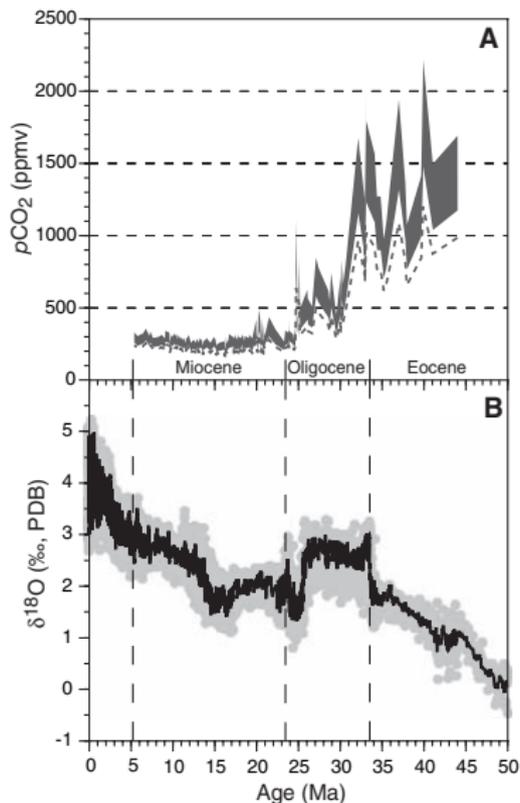
Alkenones, or $\delta^{13}\text{C}_{\text{org}}$

General approach:

Paleoatmospheric CO₂ concentrations can be estimated from the stable carbon isotopic compositions of sedimentary organic molecules known as alkenones. Alkenones are long-chained (C37-C39) unsaturated ethyl and methyl ketones produced by a few species of Haptophyte algae in the modern ocean. Alkenone-based $p\text{CO}_2$ estimates derive from records of the carbon isotopic fractionation that occurred during marine photosynthetic carbon fixation (ϵ_p). Chemostat experiments conducted under nitrate-limited conditions indicate that alkenone-based ϵ_p values ($\epsilon_{p37:2}$) vary as a function of the concentration of aqueous CO₂ (CO_{2 aq}) and specific growth rate. These experiments also provide evidence that cell geometry accounts for differences in ϵ_p among marine microalgae cultured under similar conditions.

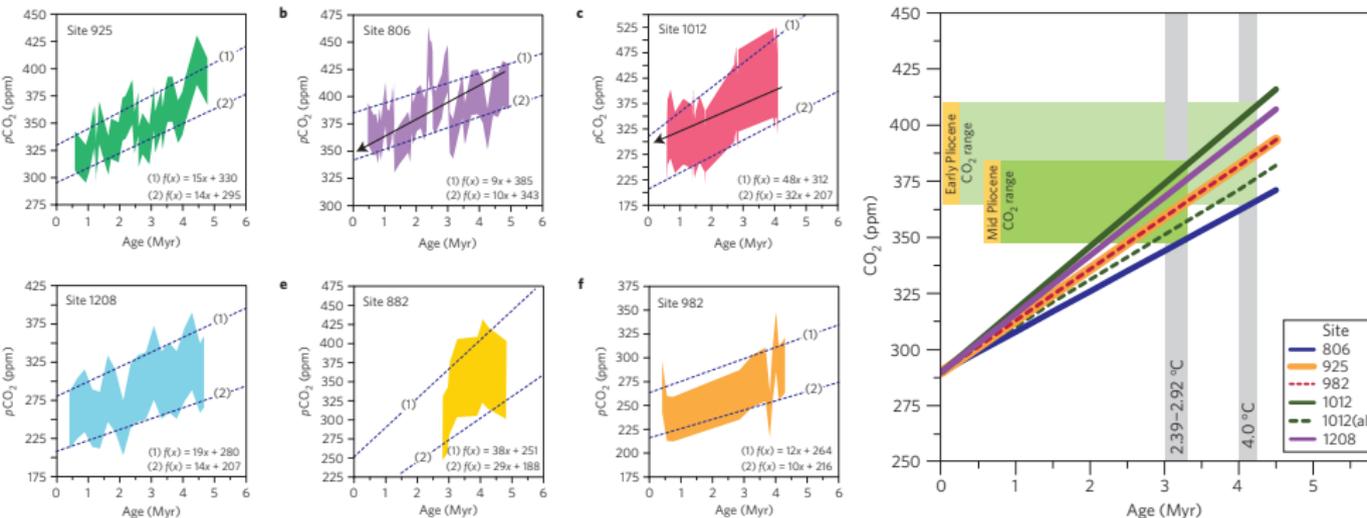
- **Disadvantage:** Based on analogue, not on chemistry, atm-surf-equilibrium might have changed over time

Alkenones, example I, last 60 Myr

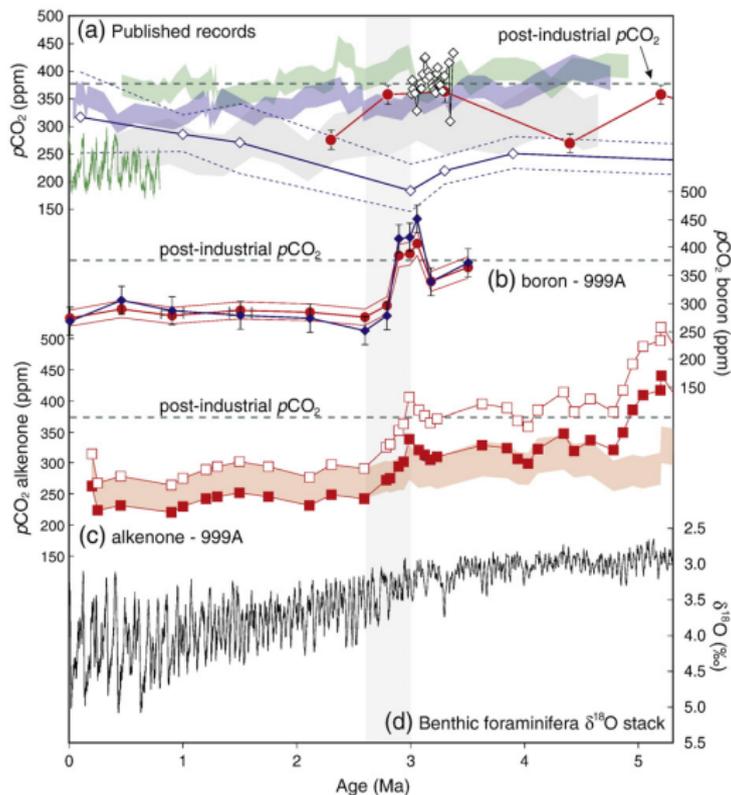


Pagani et al., 2005 S

Alkenones, example II, last 6 Myr



Pagani et al., 2010 NG

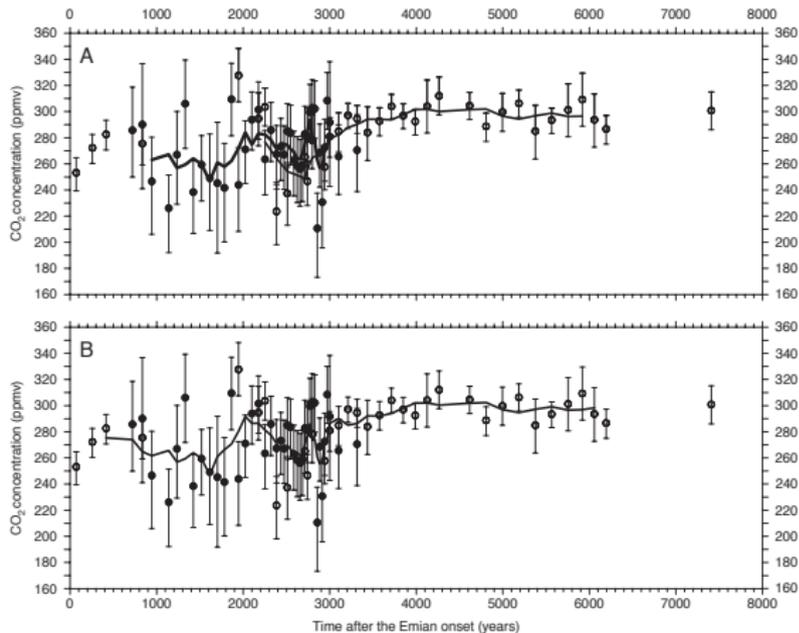
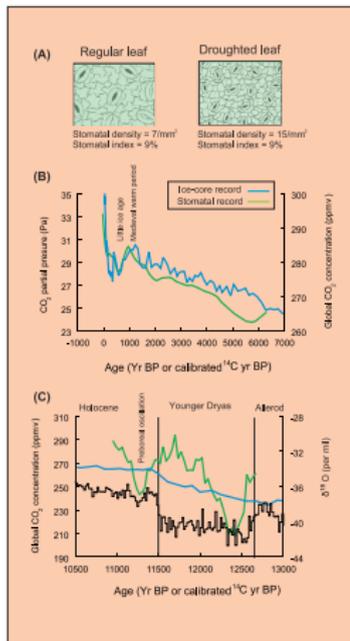
Alkenones mixed with $\delta^{11}\text{B}$, example III, last 5 Myr

Seki et al., 2010 EPSL

Outline

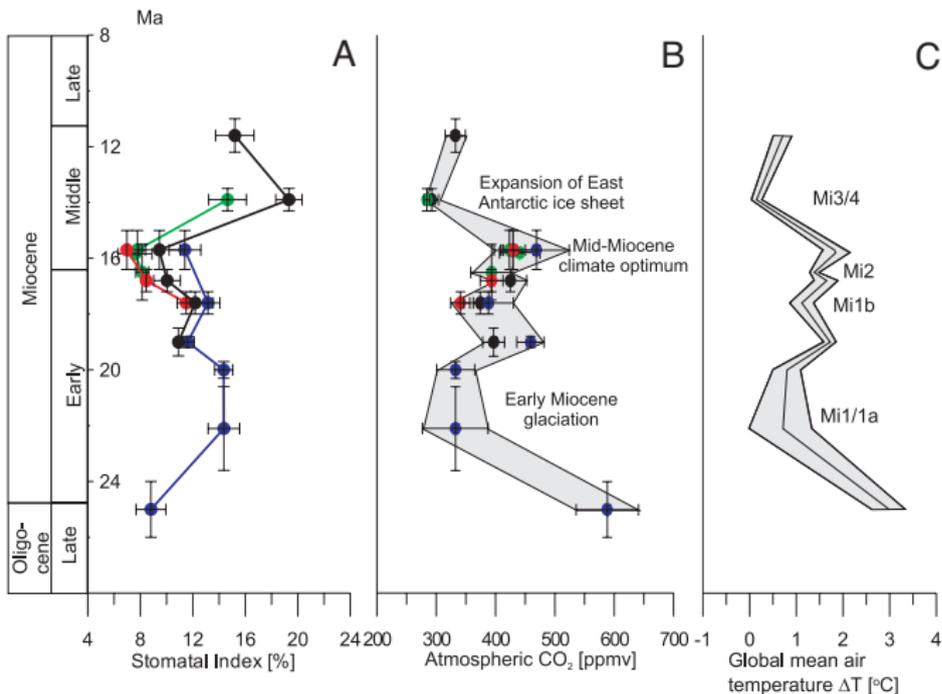
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Stomata



Rundgren 2003 GGG, Rundgren 2005 GPC

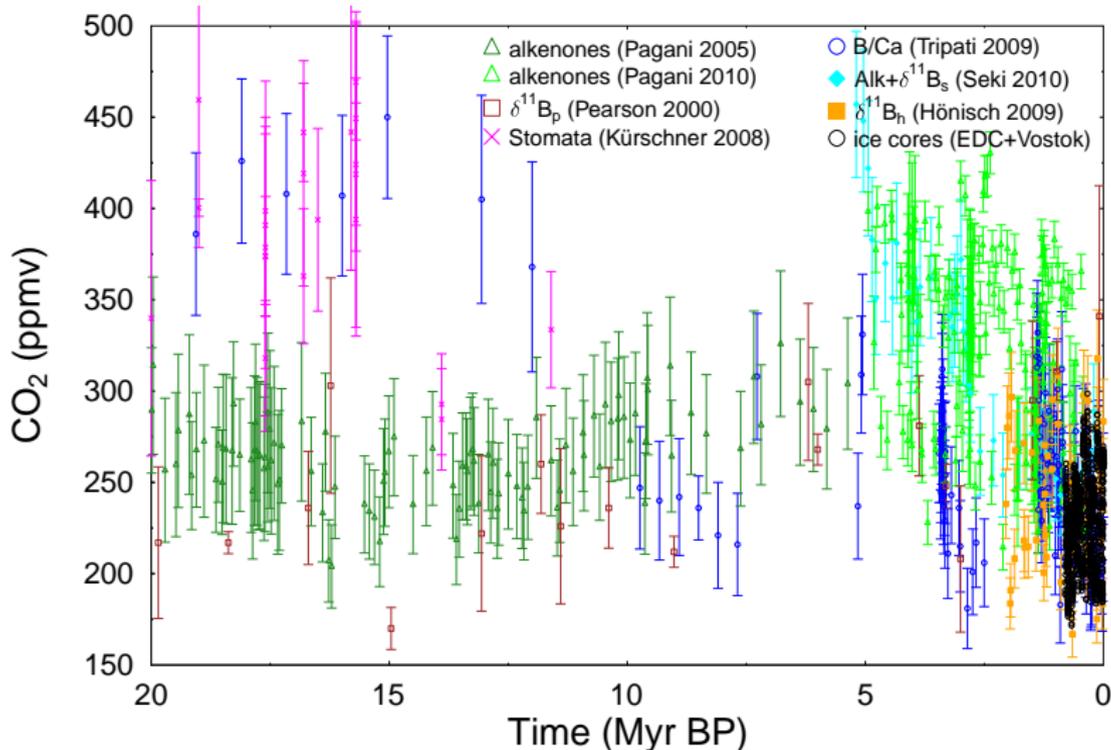
Stomata



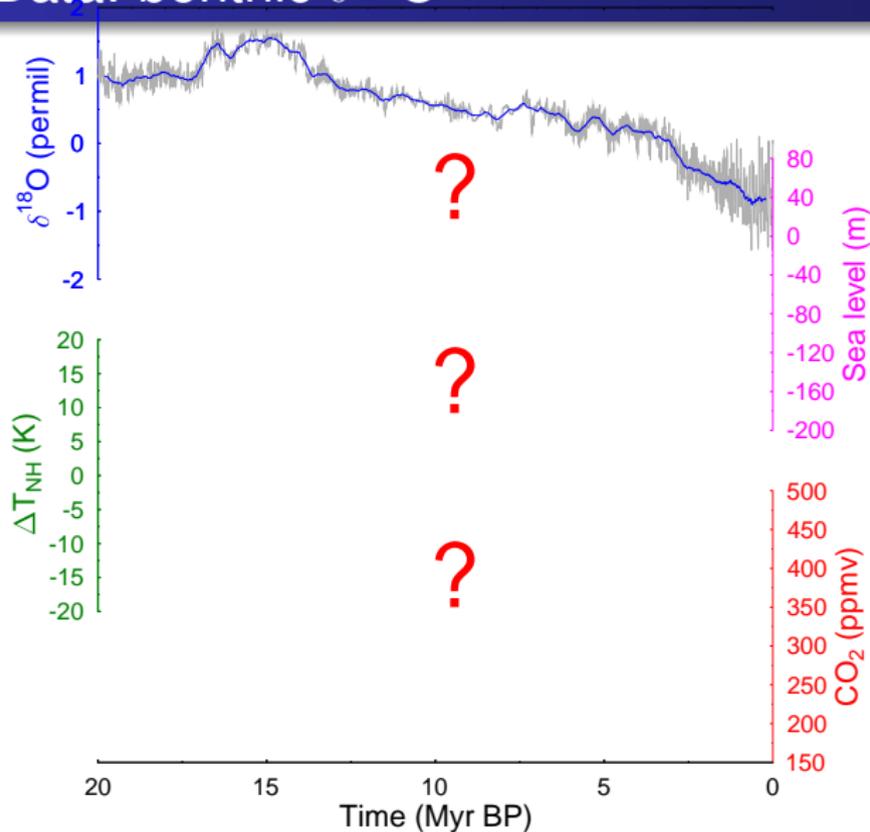
Kuerschner 2008 PNAS

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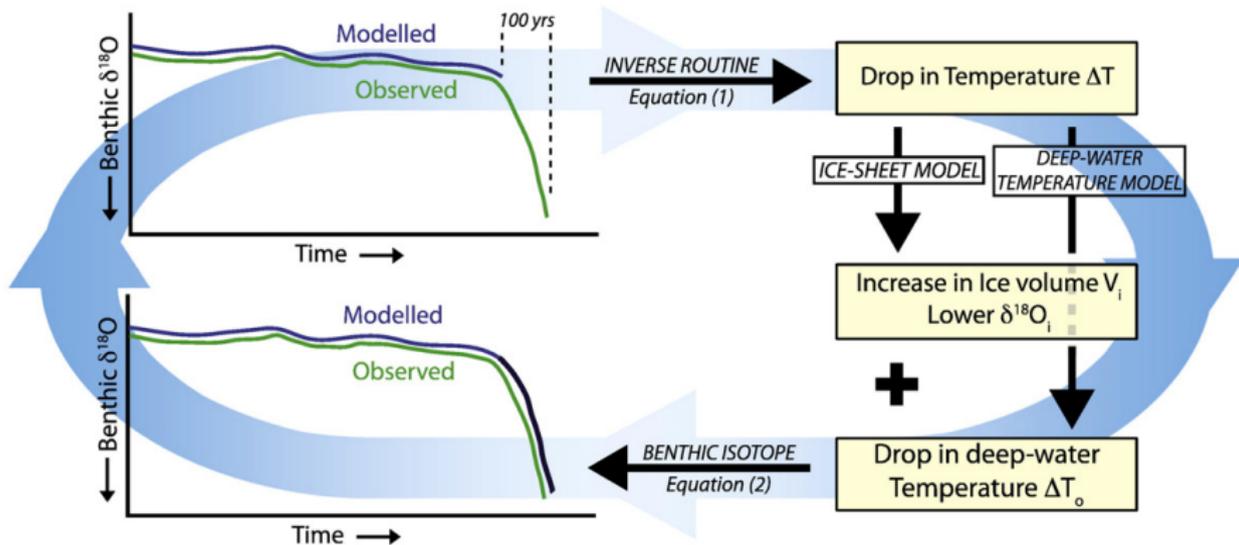
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CO₂: proxy diversity

after van de Wal et al., 2011 CPD

Climate Data: benthic $\delta^{18}\text{O}$ 

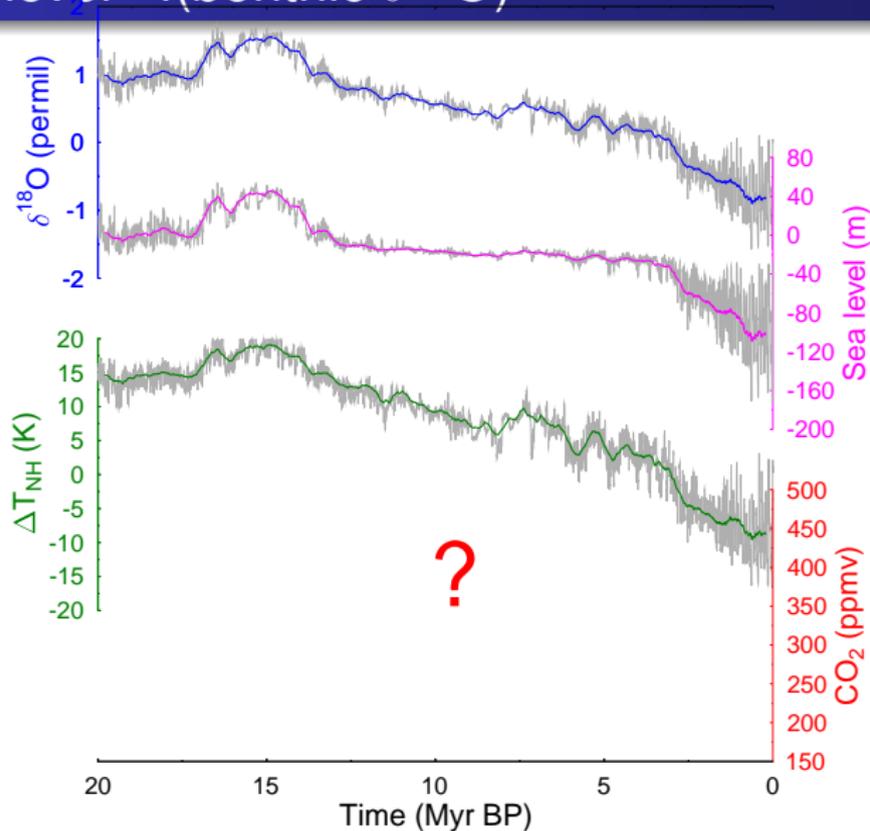
(Zachos et al., 2008, N)

Ice Sheets, ΔT and benthic $\delta^{18}\text{O}$ 

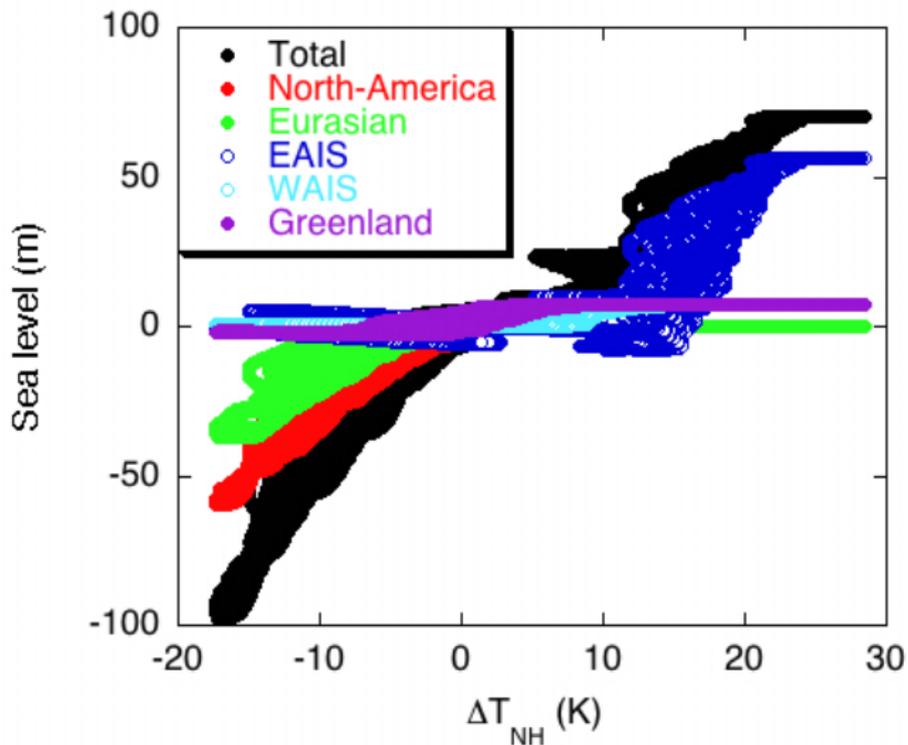
Deconvolute stacked benthic $\delta^{18}\text{O}$ into climate variables
 ($\Delta T_{\text{deep } o}$, $\Delta T_{\text{atm}} (40-80^\circ\text{N})$, size of ice sheets, sea level, snow cover)

(Bintanja et al., 2005; de Boer et al., 2011)

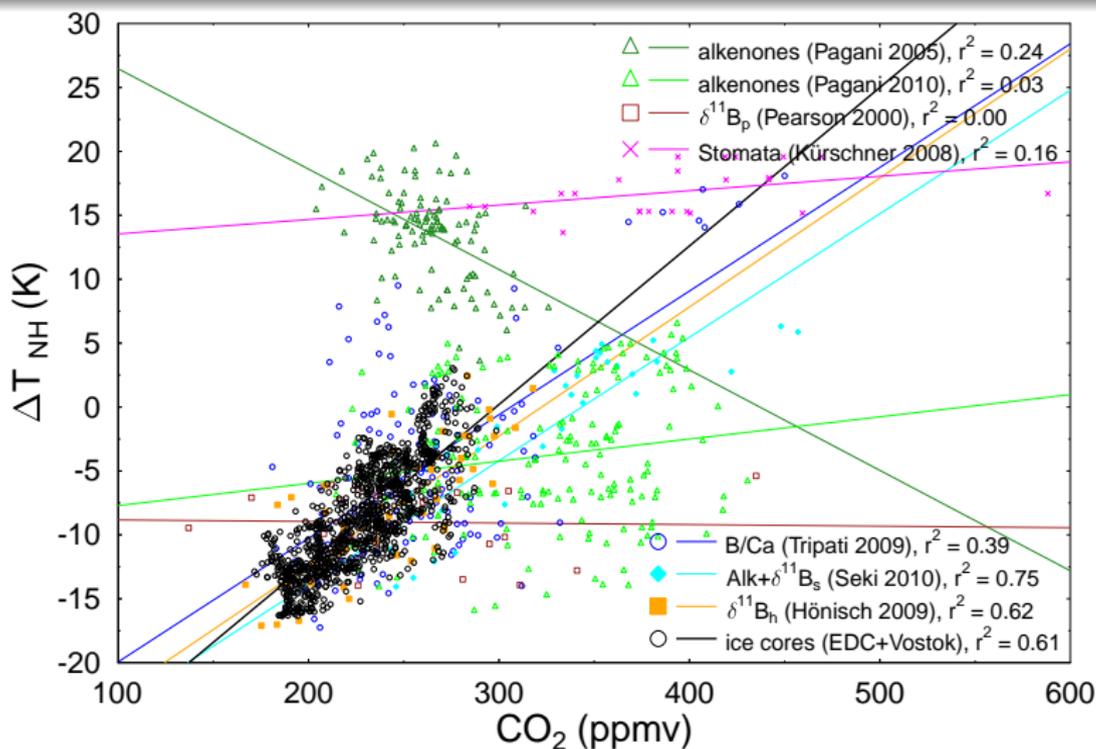
ΔT , Sea level = f(benthic $\delta^{18}\text{O}$)



(after Bintanja et al., 2005; van de Wal et al., 2011; de Boer et al., 2011)

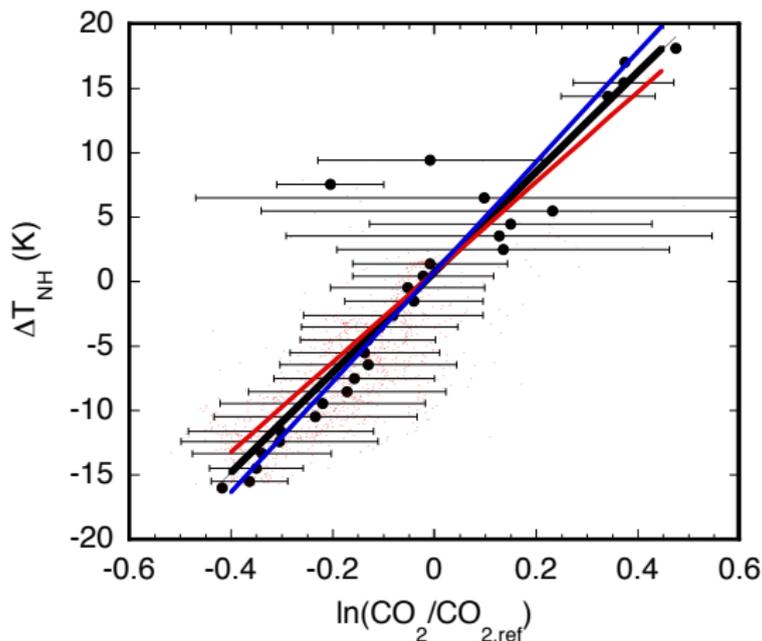
Modelling ice sheets over last 20 Myr out of $\delta^{18}\text{O}$ 

Van de Wal et al., 2011, CPD

Relationship ΔT_{NH} —CO₂

(after van de Wal et al., 2011 CPD)

ΔT_{NH} —CO₂ 1: Empirical Relationship



resampled and binned data in intervals of $\Delta(\Delta T_{NH}) = 1$ K

$C = 39 \pm 4$ K regression slope from modelled ΔT_{NH} and CO₂ data

(van de Wal et al., 2011, CPD)

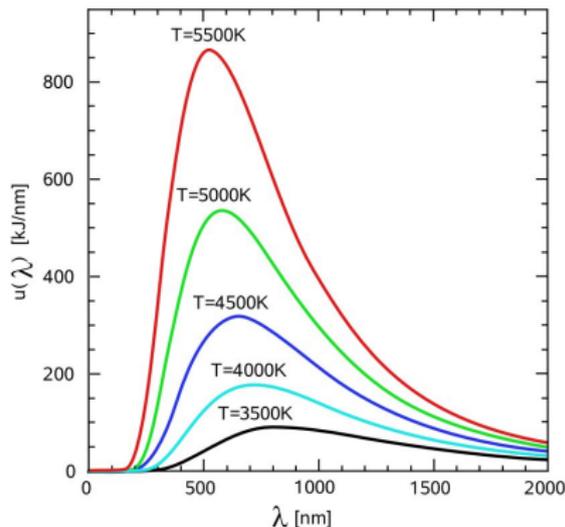
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Planck's Law

Planck's Law:
$$I(\nu, T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{\frac{h\nu}{kT}} - 1}.$$

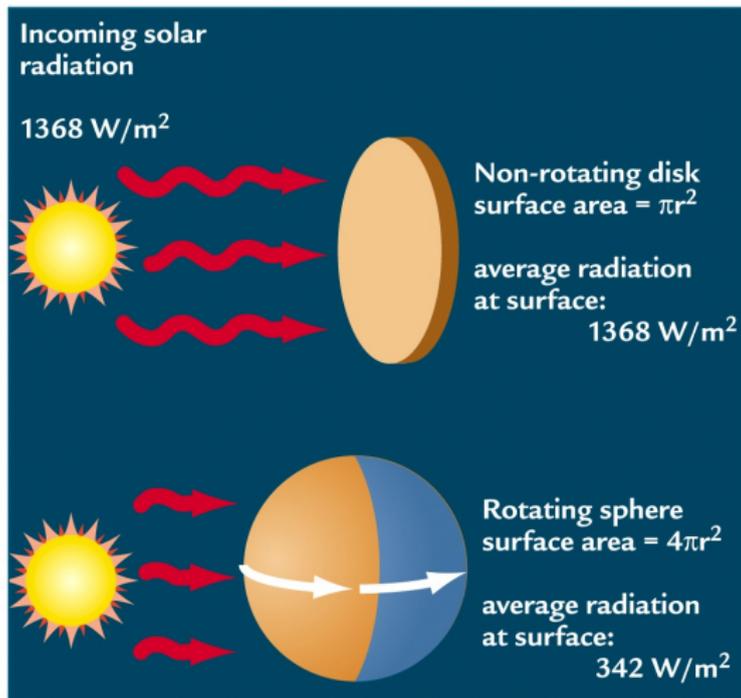
Radiation of every black body as function of temperature and wavelength.



- Birth of Quantum Mechanics: Light (photons) have discrete energies
- Planck's Constant $h \sim 6.6 \cdot 10^{-34}$ Js
- $E = h \cdot \nu$. ν : frequency
- Planck's Law brought together 2 approximations (Wien; Rayleigh-Jeans)
- Wien's displacement law:
 $\lambda_{\max} \cdot T = 2.9 \cdot 10^{-3}$ m K.
- Sun ($T = 5500$ K): $\lambda_{\max} = 527$ nm (VIS)
- Earth ($T = 255$ K): $\lambda_{\max} = 11 \mu\text{m}$ (IR)

Integration over all wavelengths: Energy emission = $f(T)$
 \Rightarrow Stefan-Boltzmann-Law: $R = \sigma T^4$

Radiation at Earth



Ruddiman 2001

Black Body Radiation

Stefan-Boltzmann-Law: $R = \sigma T^4$

Stefan-Boltzmann-Constant: $\sigma = 5.6710^{-8} \text{ W}/(\text{m}^2 \cdot \text{K}^4)$

Solarconstant: $S = 1367 \text{ W}/\text{m}^2$; average radiation: $S_M = 342 \text{ W}/\text{m}^2$.

Albedo: $\alpha = 0.3$

Steady state:

Incoming = Outgoing

$$S(1 - \alpha)\pi r^2 = R4\pi r^2$$

or

$$S_M(1 - \alpha)4\pi r^2 = R4\pi r^2$$

$$T_{e,0} = \left(\frac{S(1-\alpha)}{4\sigma} \right)^{(1/4)}$$

$$T_{e,0} = 255\text{K} (-18^\circ\text{C})$$

Measured:

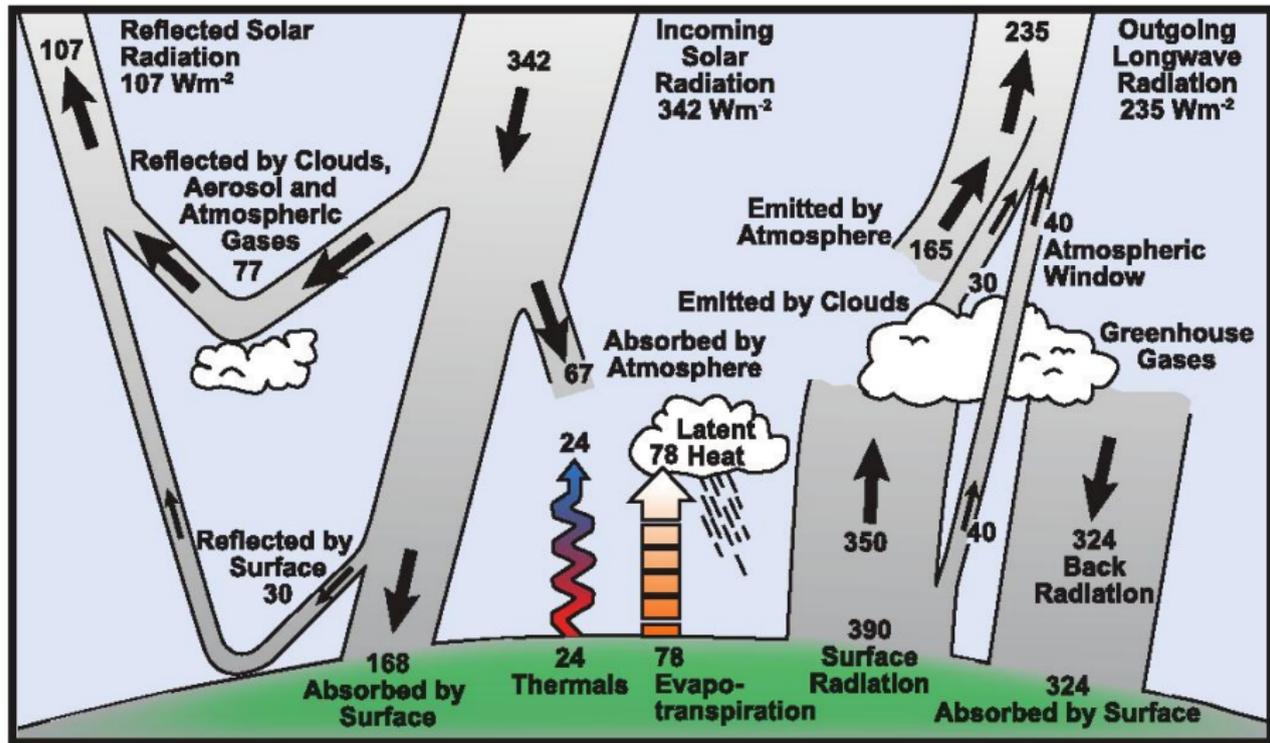
Land: 9.84°C ($1.077 \times 10^{14} \text{ m}^2$) [Leemans and Cramer(1991)]

1931–1960 Ocean: 18.1°C ($3.578 \times 10^{14} \text{ m}^2$) [Levitus and Boyer(1994)]

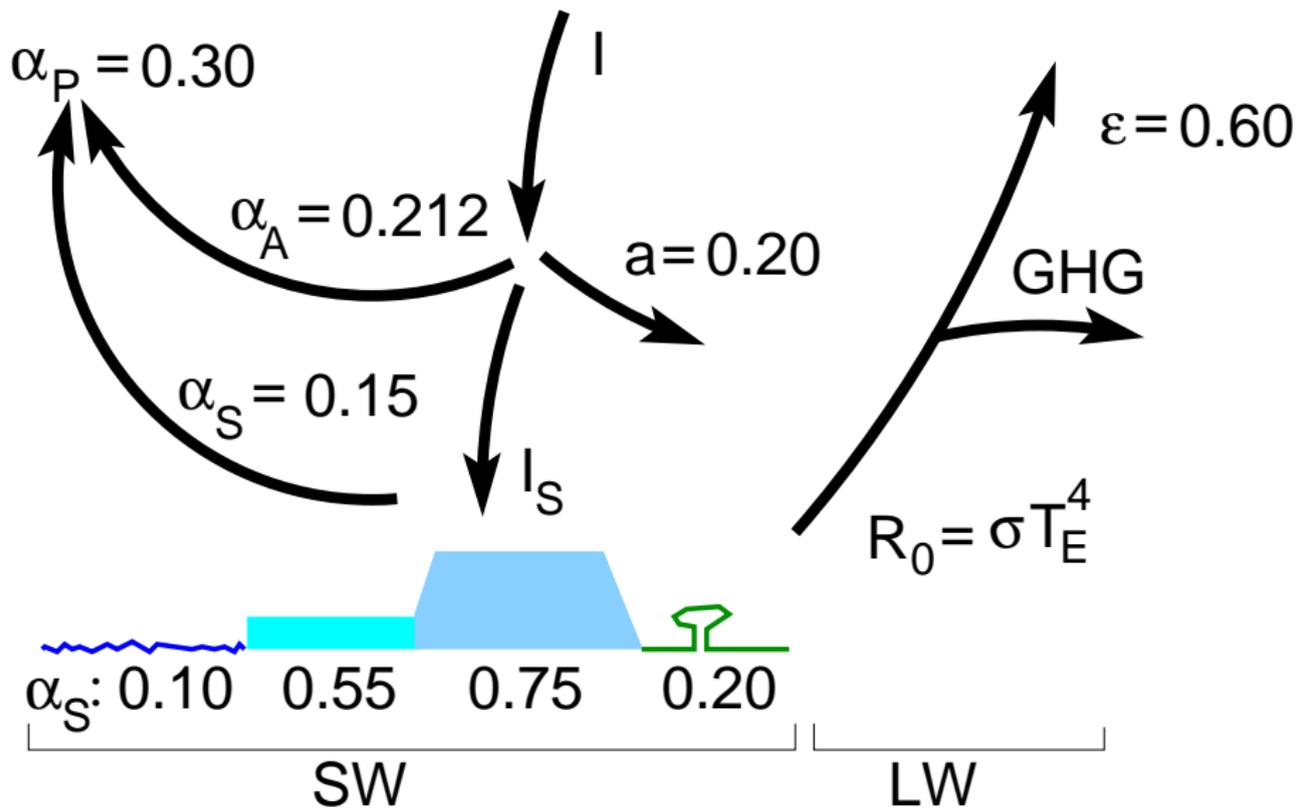
Global Mean: 16°C

Difference ($\Delta T = 34 \text{ K}$) has to be explained by radiative forcing

Energy Budget of Atmosphere (IPCC 2007)



Simplified Energy Budget (Köhler et al., 2010, QSR)



ΔT_{NH} —CO₂ 2: Theoretical Relationship

$$\Delta T_{NH} = C \cdot \ln \frac{CO_2}{CO_{2,ref}} \quad \text{with } C = \frac{\alpha\beta\gamma S_C}{1-f}$$

LGM parameters:

$$\alpha = \Delta T_{NH} / \Delta T_{global} = 15 \text{ K} / 6 \text{ K} = 2.5$$

$$\beta = 5.35 : \text{radiative forcing of CO}_2$$

$$\gamma = 1.3 : \text{enhancement factor for non-CO}_2 \text{ GHG (CH}_4, \text{ N}_2\text{O)}$$

$$S_C = 0.72 : \text{Charney climate sensitivity (fast feedbacks: Planck, water vapour, lapse rate, clouds, sea ice, albedo)}$$

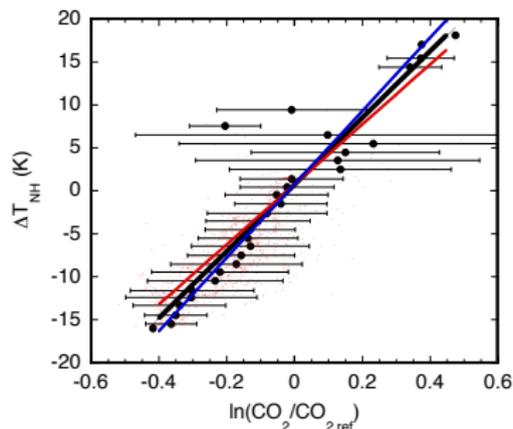
$$f = 0.72 : \text{feedbacks of slow processes (land ice, dust, vegetation)}$$

C = 43K theoretical calculation based LGM data and constant climate sensitivity

For comparison:

$$\text{pure } S_{\text{Charney}} (f = 0; \gamma = 1; \alpha = 1) \Rightarrow C_C = 3.9 \text{ K and } \Delta T_{global} = 2.7 \text{ K}$$

(van de Wal et al., 2011, CPD)

Develop relationship atmospheric ΔT_{NH} —CO₂

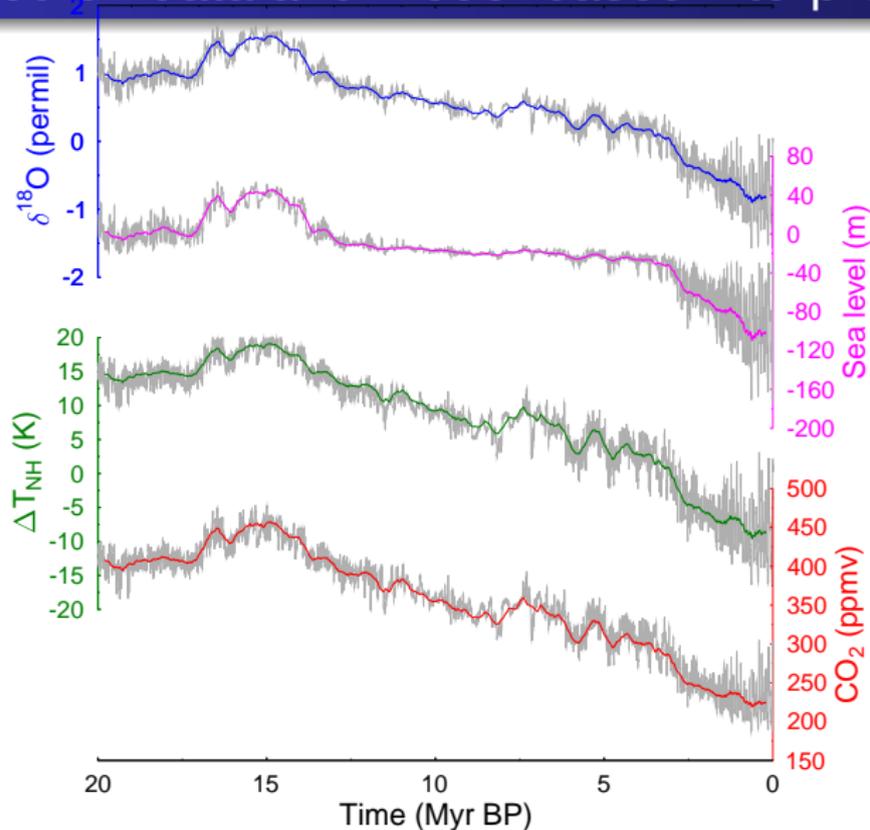
$$\Delta T_{NH} = C \cdot \ln \frac{CO_2}{CO_{2,ref}} \text{ with } C = \frac{\alpha\beta\gamma S}{1-f}$$

Two independent approaches to calculate the slope:

- 1 $C = 39 \pm 4K$ regression slope from modelled ΔT_{NH} and CO₂ data
- 2 $C = 43K$ theoretical calculation based LGM data and constant S

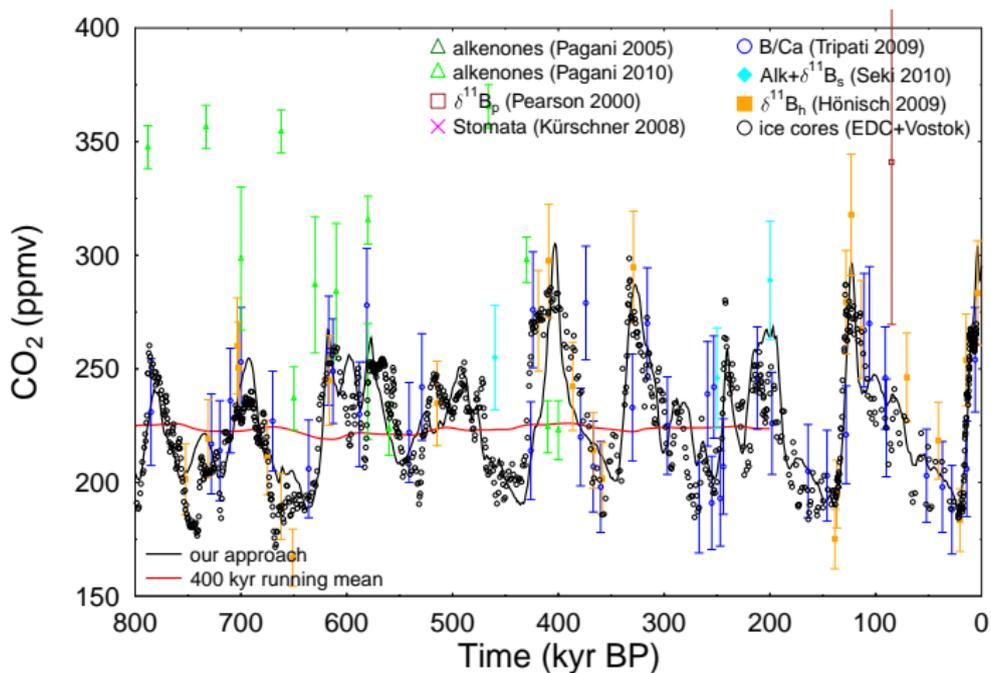
(van de Wal et al., 2011, CPD)

CO₂ based on data and model-based interpretation



after van de Wal et al., 2011 CPD

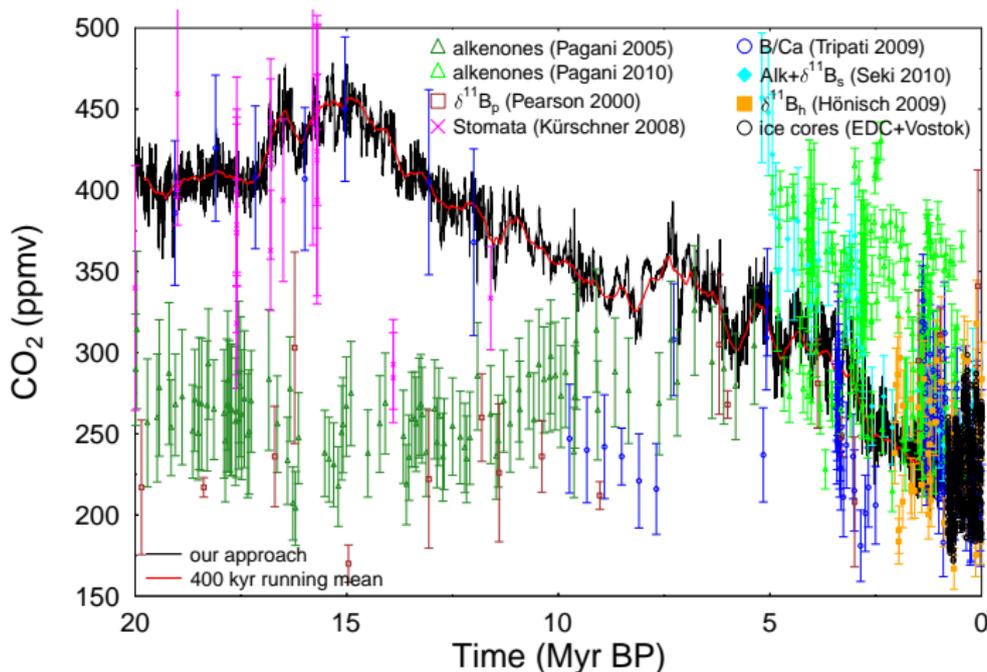
CO₂ reconstructions, the last 20 Myr



Glacial/interglacial amplitudes captured, details wrong

after van de Wal et al., 2011 CPD

CO₂ reconstructions, the last 20 Myr



Assumption: relation CO₂- ΔT unchanged with time!!!

after van de Wal et al., 2011 CPD

Validation Summary

- Calculate sea level, ΔT within one modelling framework leads to self-consistent results.
- Evaluate proxy-based CO₂ with modelling ΔT shows inconsistencies in some of the proxies (stomata, alkenones, multi-species $\delta^{11}\text{B}$)
- Regression of ΔT and best proxy-CO₂ can be understood based on theoretical background of radiative forcings
- Reconstructed CO₂ declines from 450 ppmv (20 Myr BP) to 280 ppmv at pre-industrial times.

Van de Wal et al., 2011, CPD

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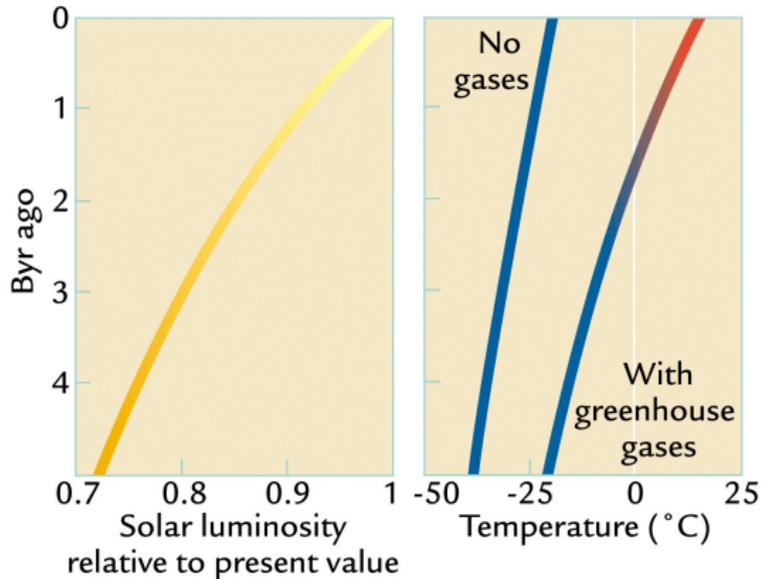
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The Faint young sun Paradox I

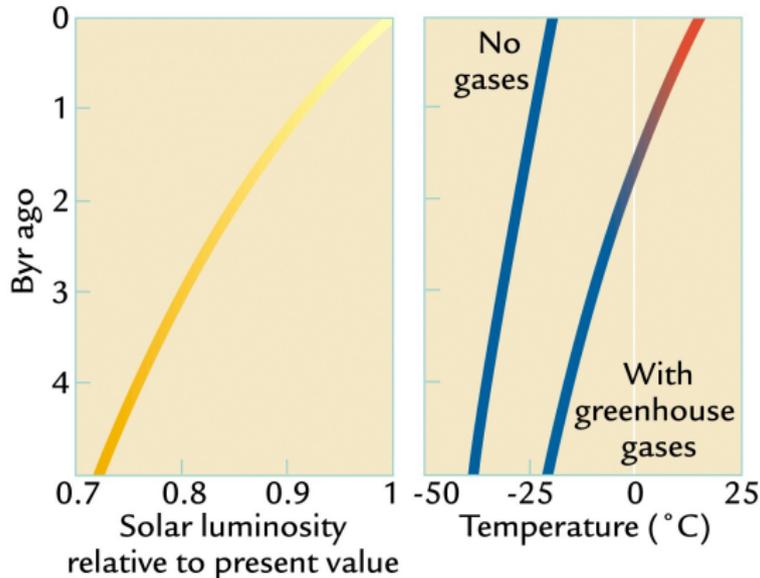
Solar luminosity increased over earth's history: Early sun was about 30% weaker than today.



At present-day atmospheric composition, temperature should have been below freezing point of water for most of earth's history

The Faint young sun Paradox I

Solar luminosity increased over earth's history: Early sun was about 30% weaker than today.



At present-day atmospheric composition, temperature should have been below freezing point of water for most of earth's history

The Faint young sun Paradox II

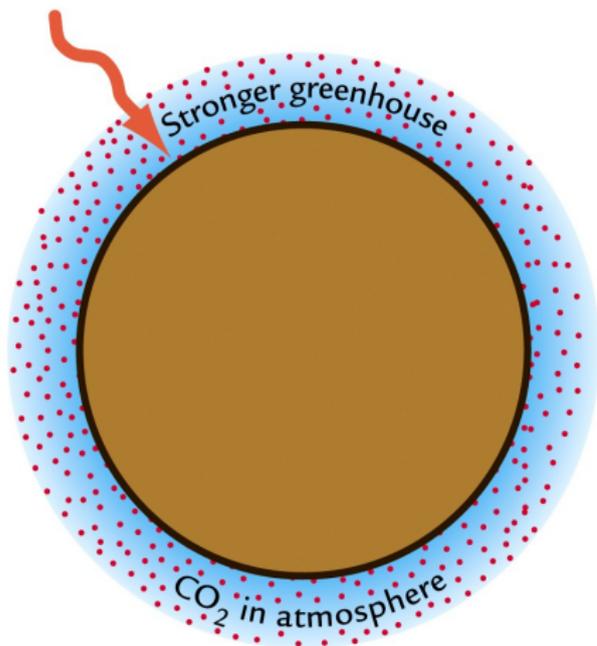
But:

- Geologic evidence for liquid ocean over at least 3.5 billion years:
Sediment rocks, microfossils showing presence of life
- Something must have prevented earth from freezing
- But if there is a heating process, it must be less active today
- Earth seems to possess a **thermostat**

Greenhouse Effect

The main candidate: A stronger greenhouse effect in early earth

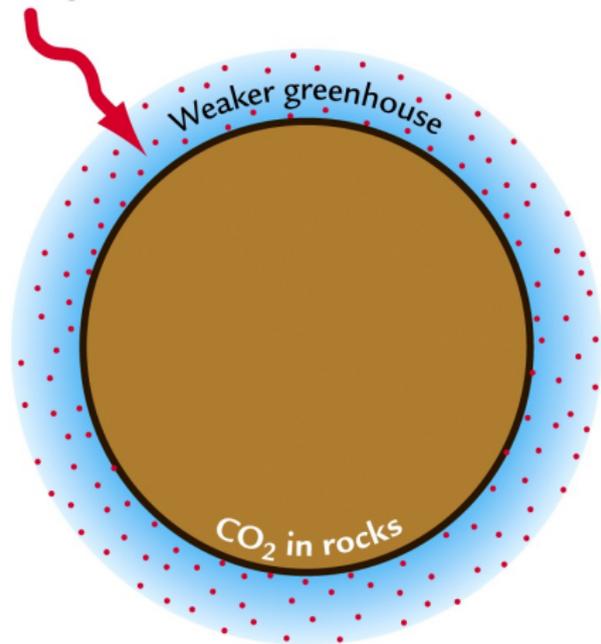
Weaker solar radiation



A

Early Earth

Stronger solar radiation

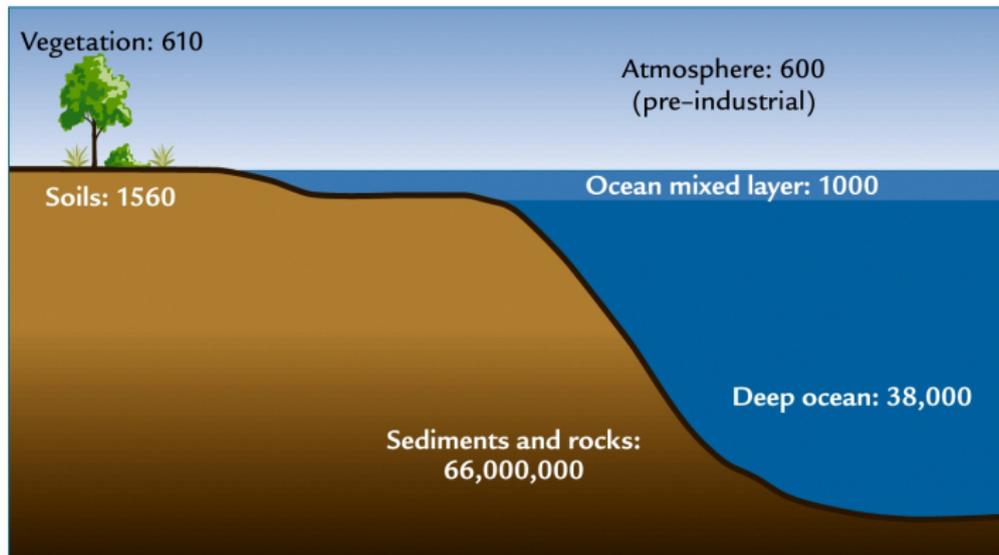


B

Modern Earth

Carbon Pools

This requires more CO₂ in the early atmosphere. Where did it come from? The largest reservoir nowadays is in rocks

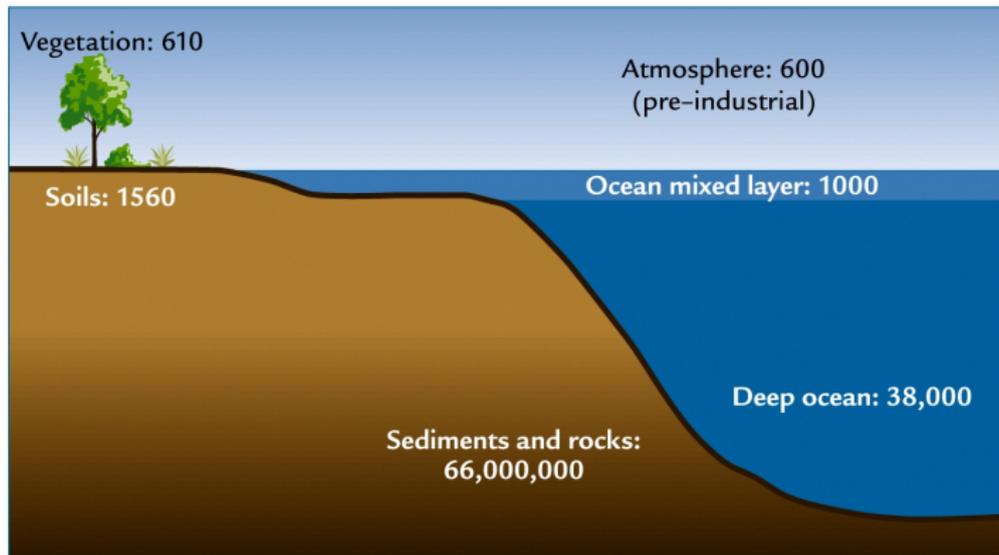


A Major carbon reservoirs (gigatons; 1 gigaton = 10^{15} grams)

How can CO₂ exchange between atmosphere and rocks?

Carbon Pools

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A Major carbon reservoirs (gigatons; 1 gigaton = 10^{15} grams)

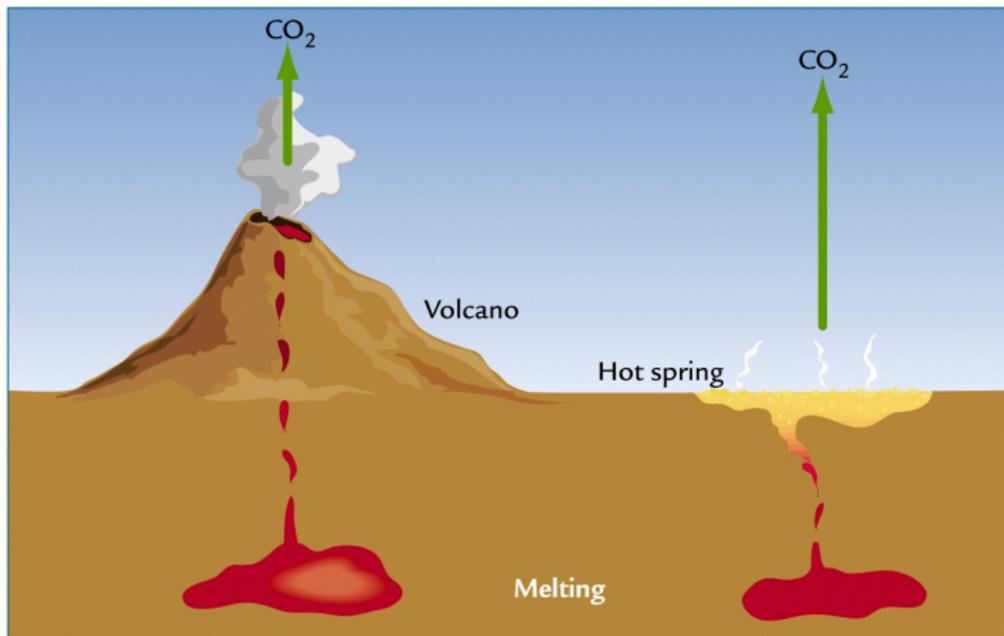
How can CO₂ exchange between atmosphere and rocks?

Outline

- 1 Basics on the Carbon Cycle
- 2 CO₂ reconstructions
 - $\delta^{11}\text{B}$
 - B/Ca
 - Alkenones, $\delta^{13}\text{C}_{\text{org}}$
 - Stomata
 - Validation of different approaches
 - Greenhouse Effect
- 3 **Processes**
 - The Faint young sun Paradox
 - **CO₂ outgassing**
 - Weathering
- 4 Summary

Rock to Atmosphere Flux: Volcanic Emissions

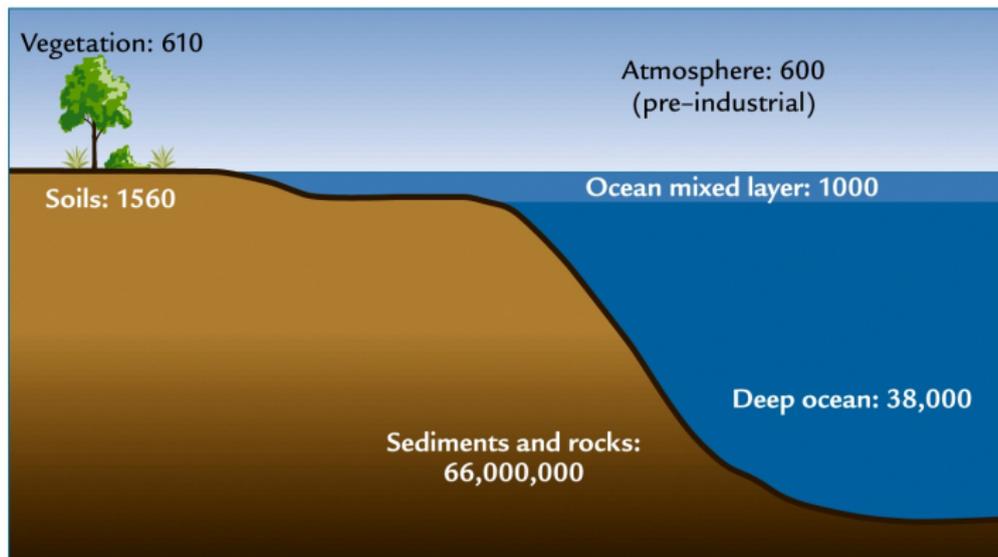
Volcanoes presently emit ca. 0.15 Pg C a⁻¹, mostly in the form of CO₂ (also some emission of CH₄). This activity might have been stronger.



Rock to Atmosphere Flux: Volcanic Emissions

Residence time of C in A/O/B with respect to volcanic outgassing:

$$\tau = \frac{41700 \text{PgC}}{0.15 \text{PgC yr}^{-1}} \approx 278000 \text{yr.}$$



A Major carbon reservoirs (gigatons; 1 gigaton = 10¹⁵ grams)

Rock to Atmosphere Flux: Volcanic Emissions

But:

- Volcanic emissions may be **drivers** of a changed CO₂ content, but they don't **react** to changes in climate.
- A thermostat requires some form of **feedback**.
- Some other process required!

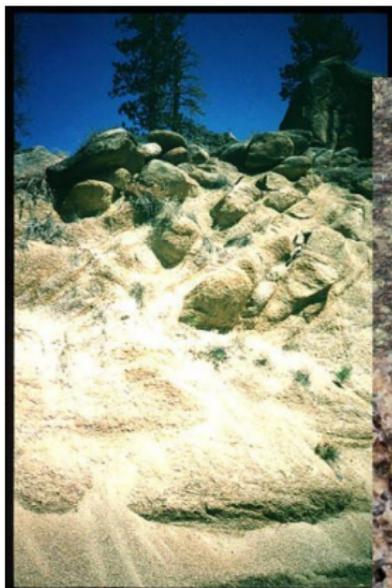
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Atmosphere to Rock Flux: Weathering

The process opposing the long-term build-up of CO_2 through volcanic outgassing is **continental weathering**.

Continental weathering is the chemical transformation of exposed rocks with rainwater and dissolved reactive gases CO_2 and O_2 .



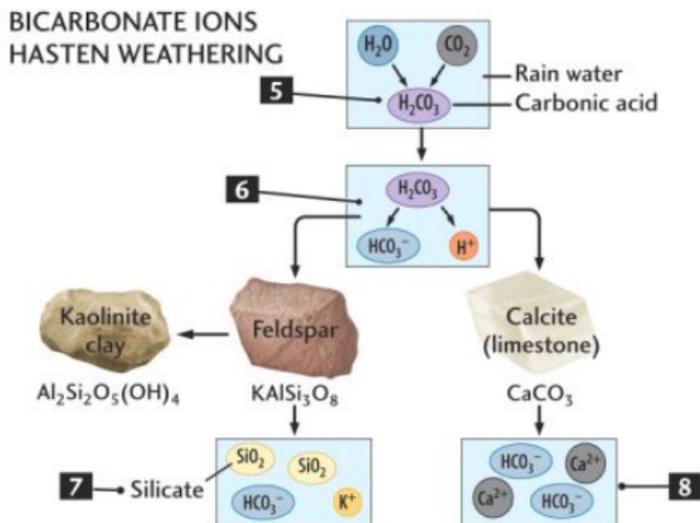
Weathered granite



Atmosphere to Rock Flux: Weathering

weathering reactions with carbonic acid in rainwater

Bicarbonate reactions



Limestone

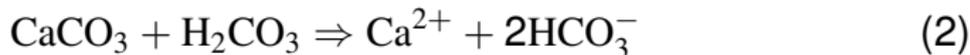


Carbonate Weathering

Limestone (CaCO_3) is easily broken down in the **dissolution** reaction



rain + atmosphere \Rightarrow carbonic acid



limestone + carbonic acid \Rightarrow continental weathering

Silicate Minerals

Typical silicate minerals: Olivine, feldspar and quartz

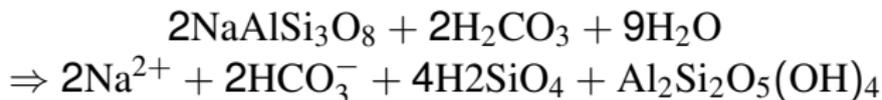


Silicate Weathering

Typical silicate weathering reaction: Na-feldspar is converted to secondary mineral kaolinite



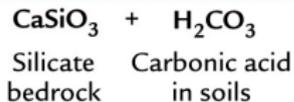
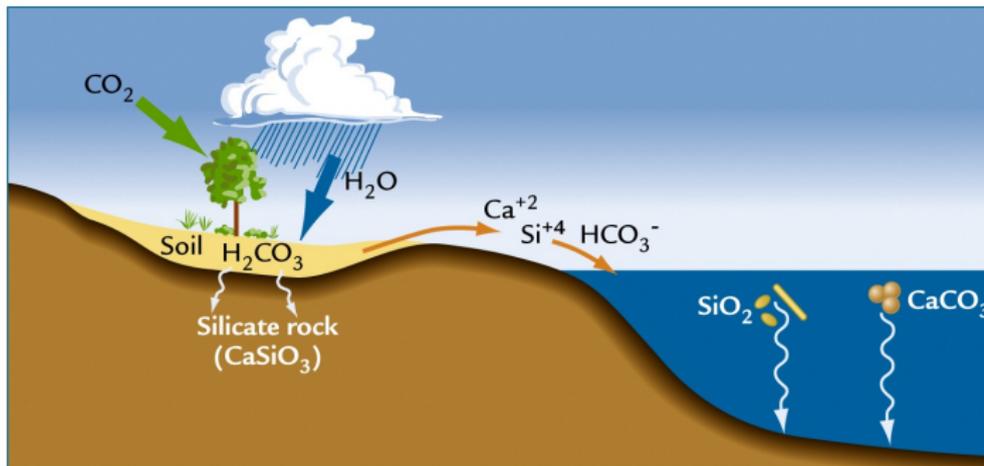
rain + atmosphere \Rightarrow carbonic acid



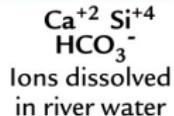
All C in silicate weathering comes from the atmosphere!

After Weathering

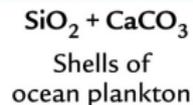
What happens with the dissolved minerals?
They are precipitated inorganically or organically.



Weathering
on land



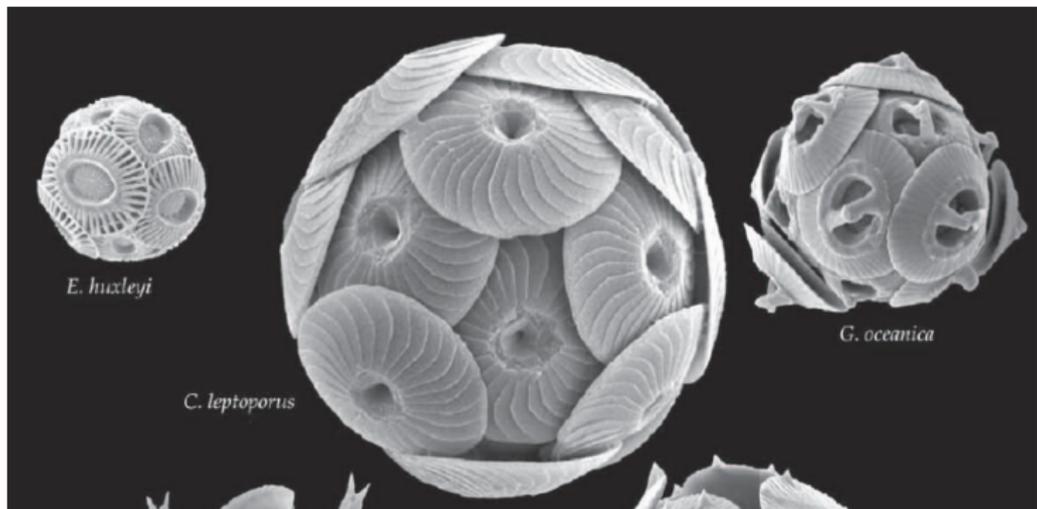
Transport
in rivers



Deposition
in ocean

Carbonate Precipitation

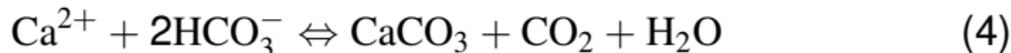
carbonate Precipitation: done by several groups, e.g. coccolithophorids



Budget of CaCO_3 pump

Organic production of CaCO_3 in the ocean:

Net reaction formula:

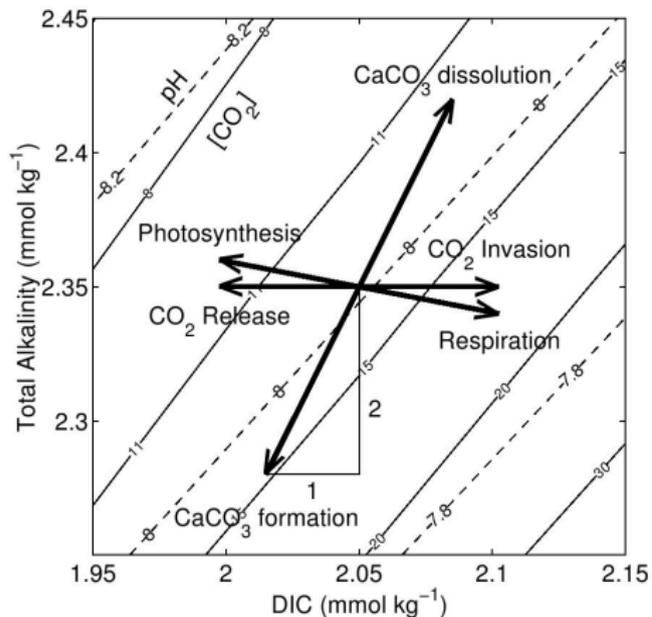


- 1 mol CaCO_3 reduced DIC by 1 mol
- 1 mol CaCO_3 reduced alkalinity by 2 mol

It is not that each mol CaCO_3 produces 1 mol CO_2 as might be suggested from this equation and the illustrations. Most of the CO_2 is immediately transformed into HCO_3^- .

However, the asynchronous changes in alkalinity and DIC change the carbonate system.

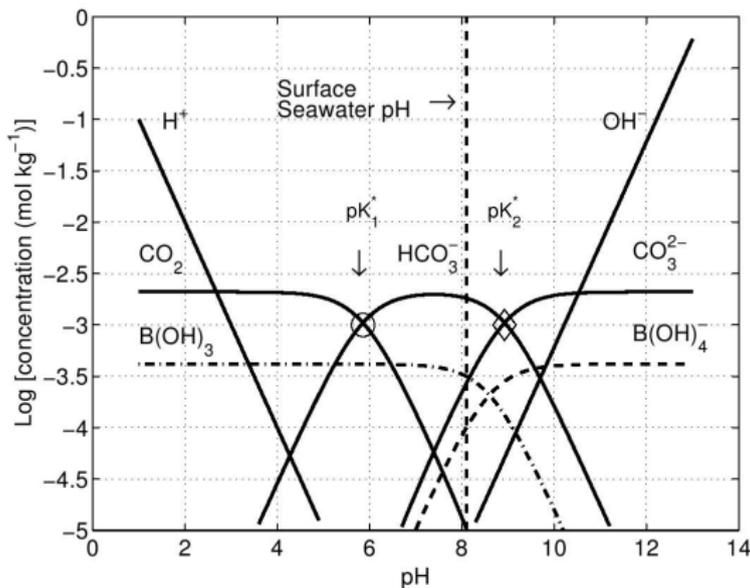
Carbonate Cycle



Zeebe & Wolf-Gladrow 2001

- CO_2 gas exchange:
 $\Delta(TA) = 0$
 \Rightarrow : CO_2 uptake reduces pH + increases $[\text{CO}_2]$
- CaCO_3 cycle:
 $\Delta(ALK) = 2 \times \Delta(DIC)$
 \Rightarrow : CaCO_3 production reduces pH + increases $[\text{CO}_2]$
- Org C cycle:
 $\Delta(ALK) = -1.14 \times \Delta(DIC)$
 \Rightarrow : Org C production increases pH + decreases $[\text{CO}_2]$

Bjerrum Plot

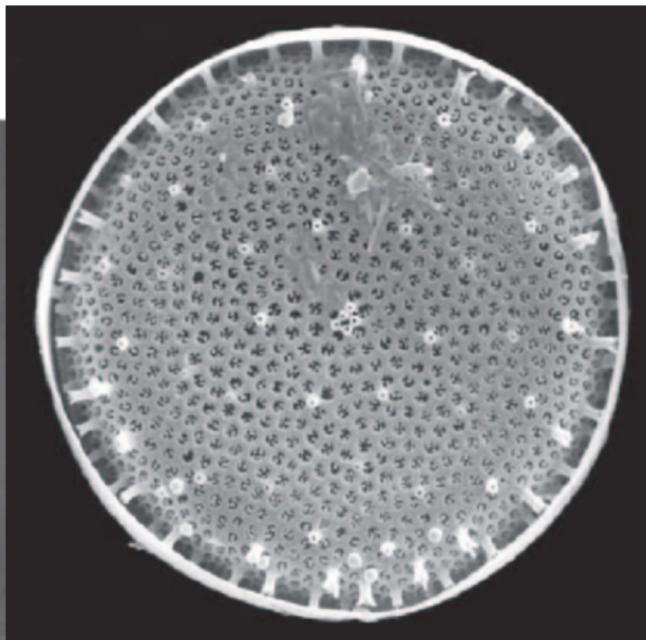
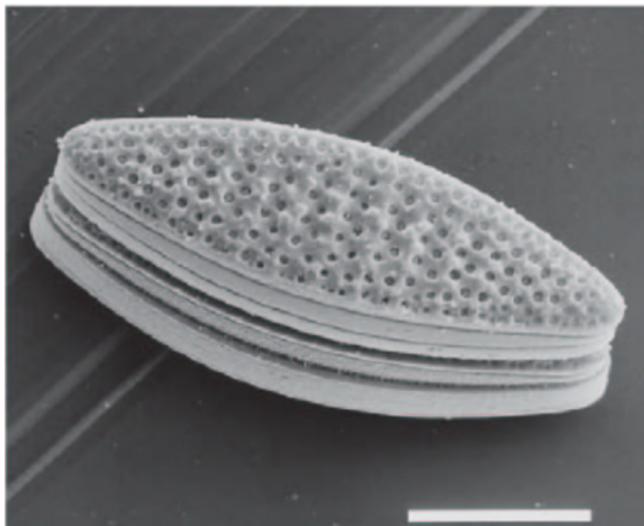


A reduced pH shifts the carbonate system towards higher CO₂ values

Zeebe & Wolf-Gladrow 2001

Silicate Precipitation

Silicate precipitation: today mostly done by diatoms



Weathering

The net effect of weathering can be summarized into the basic equation:

igneous rocks + acid volatiles \Rightarrow sedimentary rocks + salty ocean

Silicate weathering and precipitation removes CO_2 from atmosphere!

Carbonate weathering and subsequent precipitation has no net effect on CO_2 .

But both weathering processes introduce alkalinity into the ocean. So long-term effects of weathering might exist via chemical reaction of the oceanic sediment.

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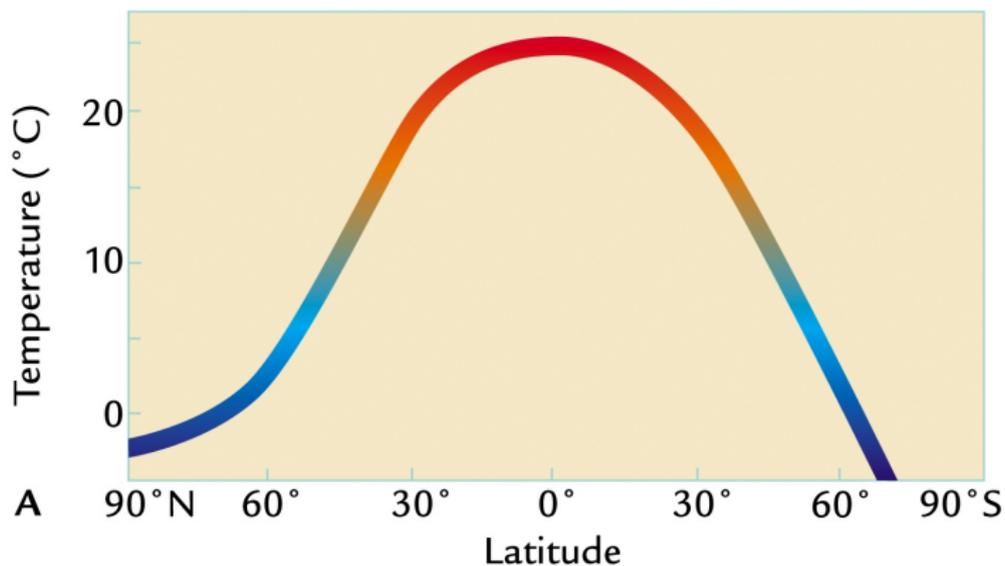
Weathering

Rate of chemical weathering depends on:

- surface to volume ratio of rock: mechanical weathering increases chemical weathering!
- temperature: reactions proceed faster in warmer climate
- precipitation: water is needed
- acidity of ground water: atmospheric CO_2 and organics have an influence

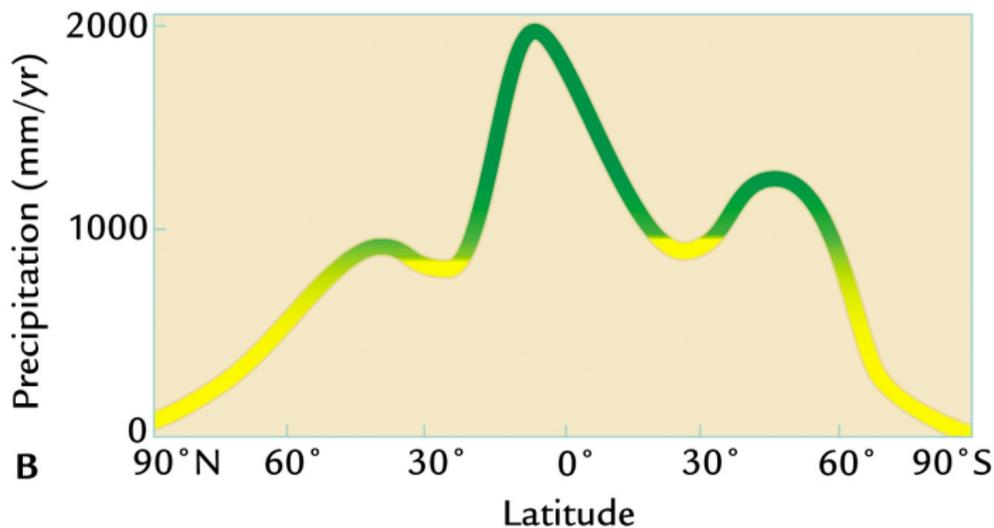
Weathering Feedback

Temperature: higher weathering in warmer regions



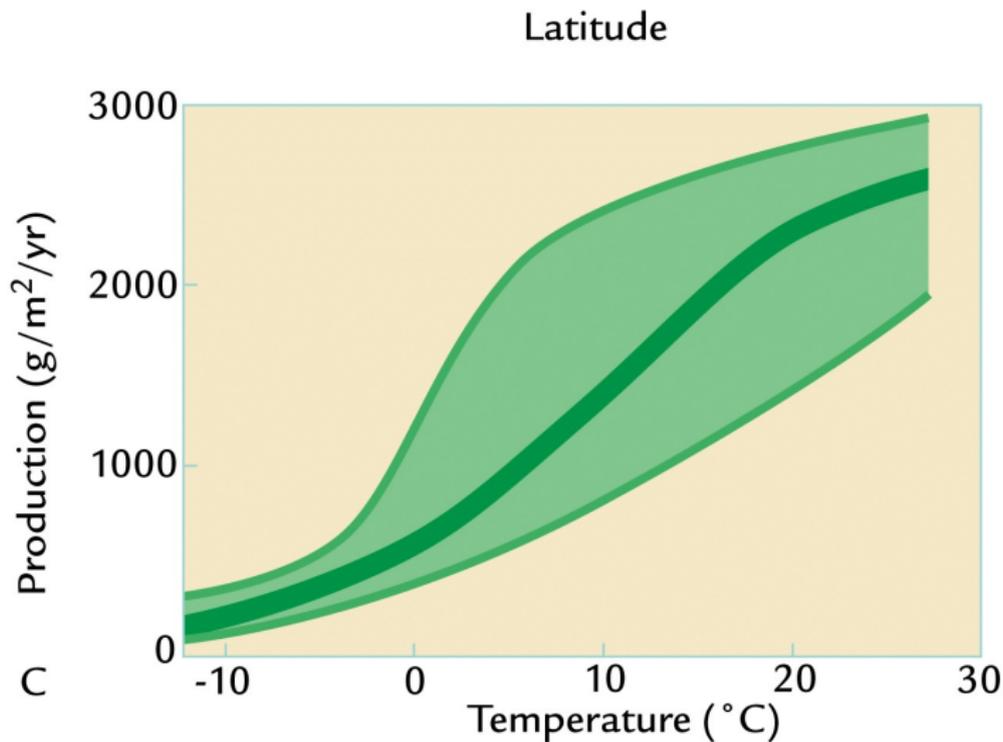
Weathering Feedback

Precipitation: highest weathering in tropics



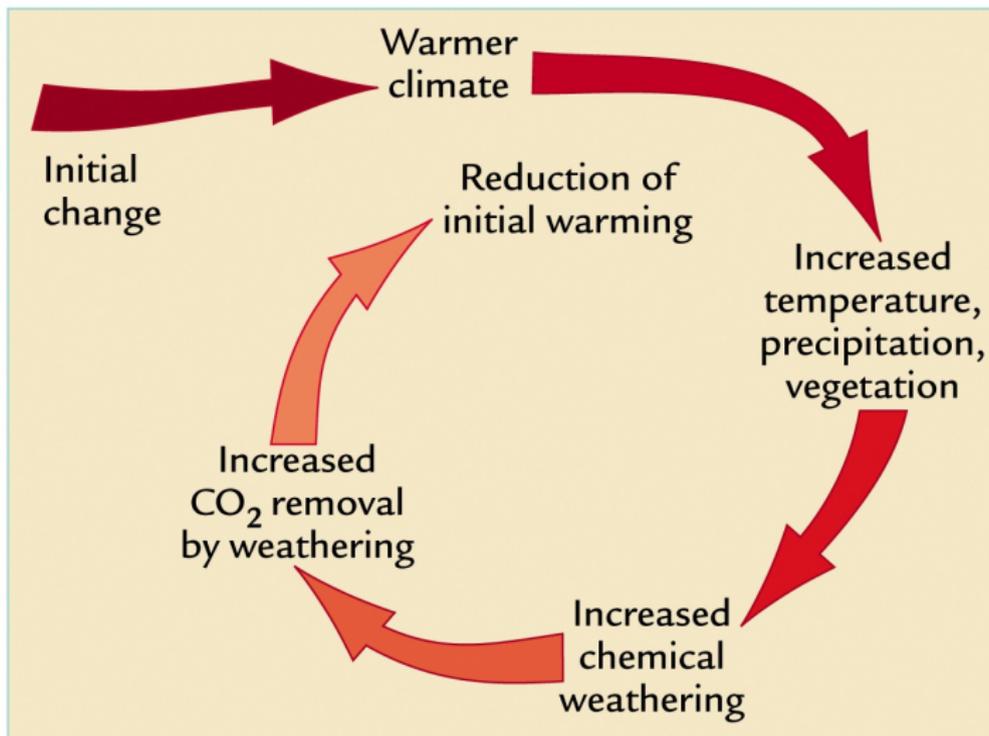
Weathering Feedback

Plant growth: increases with temperature



Weathering Feedback

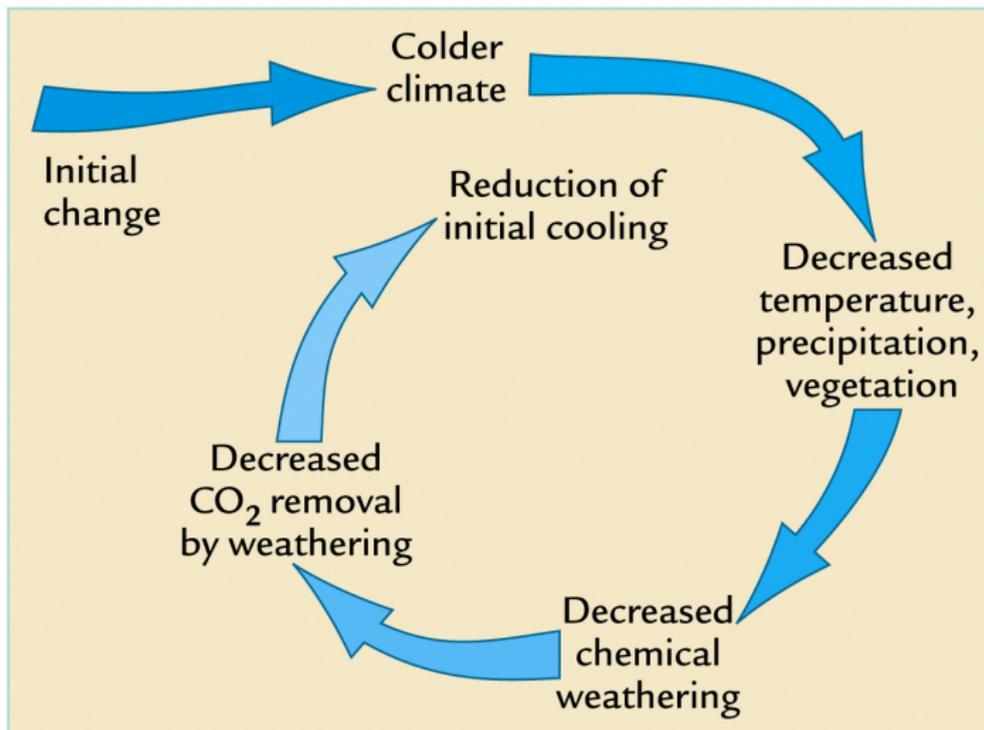
Warmer and wetter climate leads to increased weathering



A

Weathering Feedback

Sediment yield is a measure for intensity of weathering



B

Summary Weathering

Over long timescales, greenhouse strength is driven by the balance between

- source of CO_2 from volcanism
- sink of CO_2 from silicate weathering

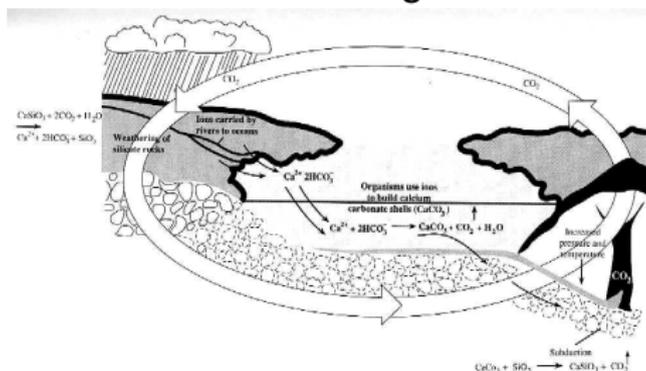


Figure 2.8 The interaction between the carbonate and silicate cycles at the surface of the Earth. Long term control of atmospheric CO_2 is achieved by dissolution of CO_2 in surface waters and its participation in the weathering of rocks. Eventually carbon is buried as part of carbonate rocks in the oceanic crust. CO_2 is released to the atmosphere when these rocks undergo metamorphism at high temperature and pressure in the Earth's crust. Modified from Keeling et al. Copyright © 1988 by Scientific American, Inc.

Important to notice:

- Changes in climate driven e.g. by CO_2 changes from volcanism.
- Negative weathering feedback dampens climate changes.
- **But that does not mean that climate does not change at all!**

Stable Cenozoic Weathering???

Vol 465 | 13 May 2010 | doi:10.1038/nature09044

nature

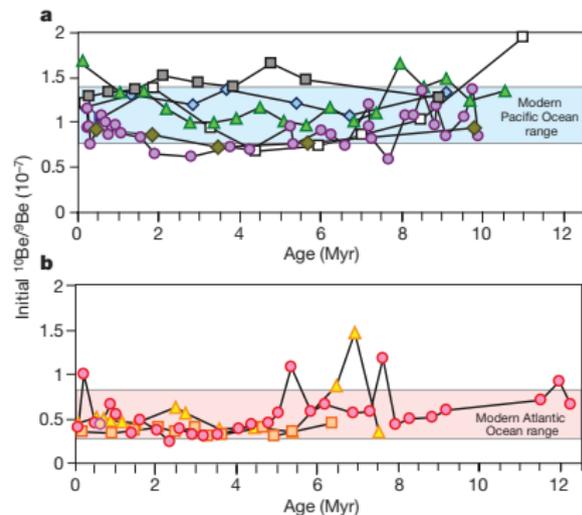
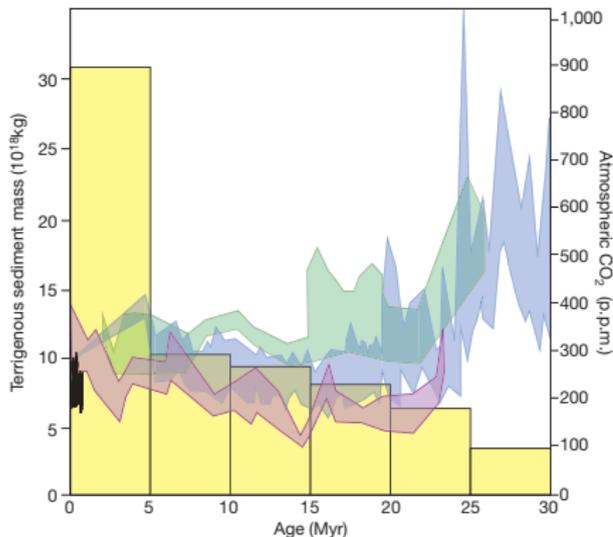
LETTERS

Long-term stability of global erosion rates and weathering during late-Cenozoic cooling

Jane K. Willenbring¹ & Friedhelm von Blanckenburg¹

Willenbring 2010 N

Stable Cenozoic Weathering???



Left: Increased sedimentation rate indicate increase in weathering
 Right: $^{10}\text{Be}/^9\text{Be}$ ratio as weathering proxy (only 10 Myr!!!)

Willenbring 2010 N

No Faint Young Sun Paradox???

nature

Vol 464 | 1 April 2010 | doi:10.1038/nature08955

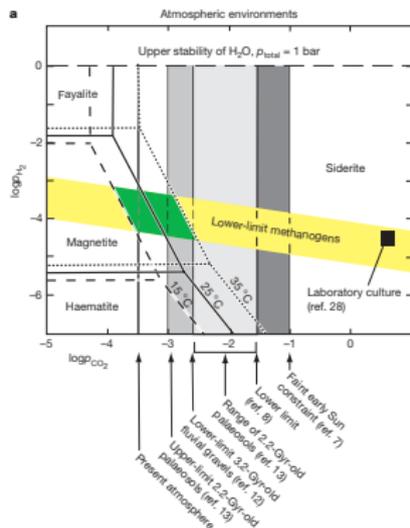
LETTERS

No climate paradox under the faint early Sun

Minik T. Rosing^{1,2,4}, Dennis K. Bird^{1,4}, Norman H. Sleep⁵ & Christian J. Bjerrum^{1,3}

Rosing 2010 N

No Faint Young Sun Paradox???



Existence of Fe(II-III) oxides (magenite) in banded iron formations is inconsitent with high CO_2 necessary under faint young sun paradox. Their solution: Lower albedo of early Earth sufficient for above freezing point.

Rosing 2010 N

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Summary

- Pre-ice core CO₂ is estimated from different proxies ($\delta^{11}\text{B}$, B/Ca, stomata, $\delta^{13}\text{C}_{\text{ORG}}$) which rather low resolution and large uncertainties.
- Validation with model-based $\Delta T = f(\delta^{18}\text{O})$ and theory on radiative forcing highlights “good” and “weak” CO₂ proxies.
- Faint Young Sun Paradox can be explained if continental weathering acts as a thermostat, which dampens climate change.
- Silicate weathering extracts CO₂ from the atmosphere and puts it in the ocean sediments.
- Carbonate weathering does not extract CO₂ from the atmosphere.
- New data weakens weathering hypothesis and Faint Young Sun Paradox.

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