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Assimilation of TOPEX/Poseidon data in a global ocean model: differences in 1995–1996

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6 Abstract

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7 Starting from an optimized climatological ocean model two response experiments are performed to study the impact of as-8 similating altimeter data on the ocean state. For this purpose TOPEX/Poseidon altimeter measurements from 1995 and 1996, re-9 spectively, are used. The model setup remains the same in the reference and the response experiments except the relative weight of 10 the altimeter data is increased for the new assimilation experiments. Furthermore a cyclic repetition of altimeter data from the respective years will be used instead of a mean annual cycle. The analysis of the differences in the two 'perpetual 1995' and 'perpetual 11 12 1996' solutions shows that the changes in the optimal forcing are small and differences in the flow fields are noticeable only in the 13 upper ocean. The model is able to follow the different sea surface height measurements by adjusting the upper ocean thermal and 14 haline structures. The modelled steric height anomalies closely follow the TOPEX/Poseidon anomalies. These anomalies are mainly 15 due to thermal expansion while the haline expansion is a second order effect in most parts of the global ocean. Nevertheless the latter 16 cannot be neglected anywhere because it at least partly compensates the thermal.

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18 Keywords: Sea surface heights; Data assimilation; TOPEX/Poseidon; Steric effect

19 1. Introduction

20 According to Gill and Niiler (1973) the variability of 21 the sea surface height (SSH) is to first order due to changes in the ocean heat content (thermal expansion) 22 23 while the effect of salinity changes plays only a second-24 ary role. However Maes (1998) and Sato et al. (2000) 25 recently demonstrated the importance of the halosteric 26 effect on SSH changes. Similar results are obtained by Levitus and Antonov (2002) who investigate the influ-27 28 ence of temperature and salinity changes on the dynamic 29 topography from measurements. A further prominent source of SSH variations is the adjustment of the ocean 30 to varying windstress fields via planetary waves (Stam-31 mer, 1997; Vivier et al., 1999). 32

The variational optimization or adjoint method provides us with a powerful tool to estimate an ocean state that is consistent with both, the given model equations and the data. This is done by optimizing certain parameters that control the models evolution in space and time: the models initial state and the surface forcing fields. Wenzel et al. (2001) (WSO hereafter) used 39 this method to estimate the climatological annual cycle 40 of the ocean. Here we will employ this method to study, 41 how it explains the interannual SSH variations as given 42 by the TOPEX/Poseidon altimeter measurements: by 43 changing the models thermo-haline structure, changing 44 the external forcing or both. 45

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2. Method and data

The purpose of this study is to obtain a global ocean 47 model state that evolves according to the model equa- 48 tions and that matches the SSH as measured by the 49 TOPEX/Poseidon altimeter mission in 1995 and 1996 50 as close as possible. As in WSO we will employ the 51 Hamburg LSG model (Maier-Reimer and Mikolajewicz, 52 1991), where LSG stands for 'Large Scale Geostrophic'. 53 This model was originally designed for ocean climate 54 studies, but has been used successfully for many other 55 purposes (e.g. Maier-Reimer et al., 1993). Here we will 56 use the coarse resolution LSG with $3.5^{\circ} \times 3.5^{\circ}$ effective 57 horizontal grid spacing and 11 layers in the vertical. The 58 great advantage of this model is the implicit formulation 59 in time, which allows for a timestep of one month. 60

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M. Wenzel, J. Schröter / Physics and Chemistry of the Earth xxx (2002) xxx-xxx

61 To combine the model and the data we use the ad-62 joint method, that is based on the ideas of Marchuk 63 (1975), LeDimet and Talagrand (1986) and others. A detailed description can be found e.g. in Thacker (1988). 64 65 The adjoint method is a variational optimization method that adjusts the models trajectory in space and 66 time to the data by optimizing certain control parame-67 68 ters while minimizing a cost function. The cost function 69 describes e.g. the squared distance between the data and 70 the corresponding model values. It may also comprise 71 constraints on the parameters themselves. A more de-72 tailed description on the implementation of the method 73 and the employed cost function can be found in WSO. 74 The final experiment therein, denoted FIN, describes the 75 models optimal climatological annual cycle as deter-76 mined by data assimilation and will serve as the first 77 guess/reference for the experiments performed here.

78 The main difference of the experiments performed in 79 this paper to the FIN experiment in WSO concern the 80 employed SSH data. While WSO use a mean annual 81 cycle from the years 1993-95, we will perform two independent experiments, TOP95 and TOP96, using the 82 83 measured annual cycles from 1995 and 1996 respec-84 tively. In both cases we will look for a cyclo-stationary 85 solution thus taking the corresponding data as perpet-86 ual, e.g. we will employ a cyclic repetition of the annual 87 cycle of the data while integrating the model for five 88 years. The SSH data are obtained from the NASA/ 89 GSFC Ocean Pathfinder Project.

90 As in WSO we will constrain the annual mean SSH 91 and the monthly anomalies seperately, where the two 92 respective annual means of the data are refered to the 93 EGM96 geoid (Lemoine et al., 1997). The climatological 94 data used are the same as in WSO. They comprise the 95 climatological annual cycle of temperature and salinity 96 from the World Ocean Atlas WOA94 (Levitus et al., 97 1994; Levitus and Boyer, 1994), estimates of the annual 98 mean transports of heat, mass and freshwater (Mac-99 donald, 1995; Slovan, 1997; Wijffels et al., 1992) as well 100 as constraints on the mean Atlantic overturning and the cyclo-stationarity of the solution. Likewise, the control 101 102 parameters that will be optimized by the adjoint method 103 are the same as in WSO: the initial model state (temperature, salinity, SSH) and the monthly forcing fields 104 105 (air temperature, surface freshwater flux, windstress).

Because we treat the data as perpetual, changes in the 106 initial state of the model will become most obvious in 107 the annual mean. Changes in the respective annual cy-108 cles will be caused mainly by different annual cycles of 109 the forcing. However from our results we cannot discern 110 the respective optimal forcings significantly. Therefore 111 we will concentrate on the interpretation of the differ-112 ences in the model's optimal annual mean states in 113 Section 3.

In contrast to WSO we will use SSH data from specific years, while the other data employed describe climatology. To obtain a model state representative for the 117 respective years the impact of the climatological datasets has to be reduced. This is done by increasing the weights 119 for altimetry by a factor of 10 as compared to WSO. This factor appeared to be a reasonable compromise to 121 achieve a good fit to the TOPEX/Poseidon data while 122 departing not too far from the given hydrography. A larger factor surely would further improve the SSH fit but at the expense of degenerating the models hydrographic state and consequently its circulation and 126 transports. 127

3. Results

Table 1 shows that the TOPEX/Poseidon data are129well reproduced in the experiments TOP95 and TOP96130respectively. Their root mean square (RMS) deviations131from the data are much less than in the reference FIN,132especially if FIN is compared to the corresponding SSH133data from 1995 and 1996. This is true for the monthly134anomalies as well as for the annual means. Conse-135quently the difference of the 1995 and 1996 annual mean136SSH is well reproduced also (Fig. 1, Table 2). The model137is able to reproduce about 70% of the spatial variance of138the difference although it slightly overestimates its am-139plitude. This is expressed by the slope of the regression140

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As compared to the reference FIN temperature and 142 salinity depart further from the WOA94 climatology 143 (Table 1). But this is intelligible because in TOP95 and 144 TOP96 the relative weights for the corresponding datasets in the cost function are reduced, which is equivalent 146 to employing larger error bars. Although the latter ex-147

Table 1

RMS	value o	of the	differences	between	the	model	solutions	and	the	corresp	onding	data
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Experiment name	Year	TOPEX/Poseidon		WOA94 (0-125 m)			
		Mean (cm)	Anomalies (cm)	Temperature (K)	Salinity (psu)		
FIN	93–95	13.2	1.27				
	95	13.3	2.02	1.21	0.283		
	96	13.1	2.11				
TOP95	95	4.64	0.38	1.98	0.347		
TOP96	96	3.88	0.28	1.95	0.340		

Details about the reference solution FIN are given in Wenzel et al. (2001).

M. Wenzel, J. Schröter / Physics and Chemistry of the Earth xxx (2002) xxx-xxx



Fig. 1. Annual mean SSH difference, $\Delta \zeta_m$ between model solutions TOP95 and TOP96. Only the area with TOPEX/Poseidon data is shown. Stippled areas exhibit negative values. Contour interval: 1 cm.

148 periments have similar RMS deviations as compared to 149 WOA94, their temperatures and salinities show up quite 150 different. Most of these differences appear in the up-151 permost 250–500 m (Fig. 2). Only in regions with deep 152 convection we also find changes in the deeper layers 153 down to the bottom.

154 In view of these differences in density (temperature 155 and salinity) and in the SSH we find only marginal 156changes in the circulation of the single model solutions. Nor there are notable differences in the forcing fields, 157 158 which might have been a candidate to explain the in-159 significant impact on the circulation. Therefore it has to be conducted that the changes in the pressure field due 160 to changes in density and SSH respectively mutually 161 162 compensate, i.e. we have to take into account the steric effect due to thermal and haline expansion. The changes 163 in SSH due to thermal and haline expansion are com-164 165 puted without any simplification as:

$$\Delta \zeta_{mT} = \int_{-H}^{\zeta} \left. \frac{1}{\alpha} \frac{\partial \alpha}{\partial T} \right|_{S,p} \Delta T \, \mathrm{d}z$$

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$$\Delta \zeta_{mS} = \int_{-H}^{\zeta} \frac{1}{\alpha} \frac{\partial \alpha}{\partial S} \bigg|_{T,p} \Delta S \, \mathrm{d} z$$

Table 2

Linear regression	between	different	annual	mean	SSH	differences	
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$x \rightarrow y$	Offset	Slope	Correlation	
$\Delta \zeta_m{}^a \to \Delta \zeta_{T/P}{}^b$	0.0314	0.9280	0.8397	
$\Delta \zeta_{mT}^{c} \rightarrow \Delta \zeta_{m}$	-0.9122	0.7648	0.8316	
$\Delta \zeta_{mS}{}^{d} \rightarrow \Delta \zeta_{m}$	-0.0584	-0.1435	-0.1090	
$\Delta\zeta_{mS} \rightarrow \Delta\zeta_{mT}$	1.0738	-0.8339	-0.5957	
$\Delta \zeta_{mS} \rightarrow \Delta \zeta_m - \Delta \zeta_{mT}$	-1.1322	0.6904	0.8290	
$\Delta\zeta_{mT} + \Delta\zeta_{mS} \rightarrow \Delta\zeta_m$	-1.1640	1.0490	0.9062	

All analysis is restricted to the region where TOPEX/Poseidon data $(\Delta \zeta_{T/P})$ are available.

^a Difference between model solutions TOP95 and TOP96.

^b Difference from 1995/96 TOPEX/Poseidon data.

^c Difference caused by thermal expansion.

^d Difference caused by haline expansion.



Fig. 2. Annual mean temperature difference from the model solutions, TOP96–TOP95, on the Atlantic section shown in the lower right inlet. Stippled areas exhibit negative values. Contour interval: 0.05 K.

respectively, where $\alpha = 1/\varrho$ is the specific volume of sea 169 water, *T*—temperature, *S*—salinity and *p*—pressure. 170 The expansion coefficients $e = (1/\alpha)(\partial \alpha/\partial T)$ and h = 171 $(1/\alpha)(\partial \alpha/\partial S)$ respectively are derived from the UNE-SCO formula for density $\varrho = \varrho(S, T, p)$ (UNESCO 173 (1981)) as it is used in the LSG model. 174

To demonstrate the concerted acting of the thermal 175 and haline expansion, $\Delta \zeta_{mT}$ and $\Delta \zeta_{mS}$ respectively, these 176 contributions to the annual mean SSH difference are 177 shown along the zonal line 11°N (Fig. 3). First of all Fig. 178 3 confirms that the differences $\Delta \zeta_{T/P}$ as given by the 179 TOPEX/Poseidon data are well reproduced by the 180 models difference $\Delta \zeta_m$, TOP96–TOP95. Additionally we 181 find different situations for the varying role of $\Delta \zeta_{mT}$ and 182 $\Delta \zeta_{mS}$. While most of $\Delta \zeta_m$ can already be explained by 183 ad ζ_{mT} in the western part of the Indian Ocean and in the 184 Pacific, $\Delta \zeta_{mT}$ and $\Delta \zeta_{mS}$ have the same magnitude but 185 opposit sign in the eastern part of the Indian Ocean and 186 in the Atlantic. Here they nearly cancel. 187

In total $\Delta \zeta_{mT}$ (Fig. 4a) is capable of reproducing 188 about 69% of the spatial variance in $\Delta \zeta_m$ (Table 2), i.e. 189 the correlation amounts to 0.83, but the amplitude of 190 the thermal expansion appears too big as indicated by 191 the slope of the regression which is only 0.76. This 192 overshooting of the thermal expansion is compensated 193 by the halosteric effect, $\Delta \zeta_{mS}$ (Fig. 4b), whose sign usu- 194 ally is opposite to the thermal at least in most parts of 195 the world ocean and whose horizontal structure is sim- 196 ilar to that of $\Delta \zeta_{mT}$, thus leading to a negative correla- 197 tion (Table 2). Both fields show values ranging from -5 198 to +8 cm. Maximum values in $\Delta \zeta_{mT}$ we find in the At- 199 lantic and in the western tropical Pacific as well as in the 200 eastern part of the Indian Ocean (Fig. 4a). For $\Delta \zeta_{mS}$ we 201 find the largest positive value at the boundaries and in 202 the region south of Africa. The largest negative halos- 203 teric changes are located in the southern Indian Ocean 204 M. Wenzel, J. Schröter | Physics and Chemistry of the Earth xxx (2002) xxx-xxx



Fig. 3. Annual mean SSH difference, $\Delta \zeta_m$ (solid line with asterisk), from the model solutions, TOP96–TOP95, along the zonal section 11°N as compared to TOPEX/Poseidon, $\Delta \zeta_{T/P}$ (thick solid line). $\Delta \zeta_m$ is split into the parts caused by thermal and by haline expansion, $\Delta \zeta_{mT}$ (dashed with triangles) and $\Delta \zeta_{mS}$ (dashed with dots) respectively, as well as the residual, $\Delta \zeta_m - \Delta \zeta_{mT} - \Delta \zeta_{mS}$ (solid with circles) which is due to changes in the circulation.



Fig. 4. Annual mean SSH difference between model solutions TOP95 and TOP96 caused (a) by the thermosteric effect, $\Delta \zeta_{mT}$, and (b) by the halosteric effect, $\Delta \zeta_{mS}$. Stippled areas exhibit negative values. Contour interval: 1 cm. As in Fig. 1 only the area with TOPEX/Poseidon data is shown.

205 and especially in the tropical North Atlantic where it 206 nearly compensates the thermosteric.

207 Although $\Delta \zeta_{mS}$ has only low correlation to the total 208 SSH differences $\Delta \zeta_m$ (-0.11) it reflects the difference



Fig. 5. Relative magnitude of the thermal and the haline expansion in percent. High values (hatched, above 60) indicate that the main contribution to the SSH changes stem from thermal expansion, while low values (stippled, below 40) indicate the dominance of the halosteric effect. In areas with no special signature thermosteric and halosteric effect nearly compensate each other because of their opposite sign. Contour interval: 20. As in Fig. 1 only the area with TOPEX/Poseidon data is shown.

 $\Delta \zeta_m - \Delta \zeta_{mT}$ well. Consequently the combined thermo- 209 haline expansion explains the SSH differences best. The 210 correlation is improved to be 0.91 now, i.e. $\Delta \zeta_{mT} + \Delta \zeta_{mS}$ 211 is responsible for about 83% of the spatial variance in 212 $\Delta \zeta_m$. Furthermore the slope of the regression is im- 213 proved from 0.76 considering thermal expansion only to 214 1.05 including the halosteric effect. We only partly agree 215 with Gill and Niller (1973) that the latter plays a minor 216 role. Concerning the magnitude this is true in many 217 parts of the global ocean. But there are also large re- 218 gions, especially the south-east Pacific and much of the 219 Atlantic, where $\Delta \zeta_{mS}$ exceeds the thermosteric effect or at 220 least has a comparable magnitude (Fig. 5). Furthermore 221 the thermosteric and the halosteric effect are anticorre- 222 lated (Table 2), thus compensating each other. Therefore 223 the halosteric effect should not be neglected anywhere! 224 Similar results are obtained e.g. by Maes (1998), who 225 investigates the results of different model configurations, 226 or by Levitus and Antonov (2002), who compute the 227 influence of temperature and salinity changes on the 228 dynamic topography from measurements. Furthermore 229 Sato et al. (2000) demonstrated the improvents achieved 230 in estimating the heat storage changes from TOPEX/ 231 Poseidon when additionally considering salinity chan-232 ges. Finally, only a minor SSH residual, $\Delta \zeta_m - \Delta \zeta_{mT} -$ 233 $\Delta \zeta_{mS}$, remains that contributes to changes of the pres- 234 sure field (Fig. 6) and because of its weak gradients it 235 has only little influence on the circulation. 236

4. Summary and conclusions

This paper demonstrates the capability of the LSG 238 model to reproduce the differences in the annual mean 239 SSH for 1995 and 1996 as seen by the TOPEX/Poseidon 240

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M. Wenzel, J. Schröter / Physics and Chemistry of the Earth xxx (2002) xxx-xxx



Fig. 6. The residual part of the models SSH difference, $\Delta \zeta_m - \Delta \zeta_{mT} - \Delta \zeta_{mT} - \Delta \zeta_{mT}$, that contributes to pressure field changes at depth. The area mean value as given by the offset in Table 2 (last row) has been removed because it has no bearing on the circulation. Stippled areas exhibit negative values. Contour interval: 0.5 cm. As in Fig. 1 only the area with TOPEX/Poseidon data is shown.

241 altimeter. It appears that the resulting SSH differences 242 have only little influence on the circulation. A detailed 243 inspection of the results showed that the optimization 244 method prefers to change the initial thermo-haline 245 structure, i.e. the density field of the model, instead of 246 changing the forcing fields to achieve a good cyclo-sta-247 tionary solution. Most of these changes in density ap-248 pear in the uppermost 250–500 m of the water column. 249 Only in regions with deep convection we also find 250 changes in the deeper layers down to the bottom. They 251 compensate the SSH differences, thus there are only 252 minor differences in the pressure field which is the main 253 defining quantity for the velocities. Due to the steric 254 effect the changed thermo-haline structure is responsible 255 for about 82% of the modelled SSH differences. 69% of 256 the total can already be explained by thermal expansion 257 which is the dominant effect in most of the global ocean. 258 The haline expansion usually is weaker than the thermal one, but because they are anti-correlated they partly 259 260 compensate. Therefore one should not neglect the 261 halosteric effect when interpreting SSH differences.

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